



# E-TEST prototype and advanced GW technologies

Chiara Di Fronzo

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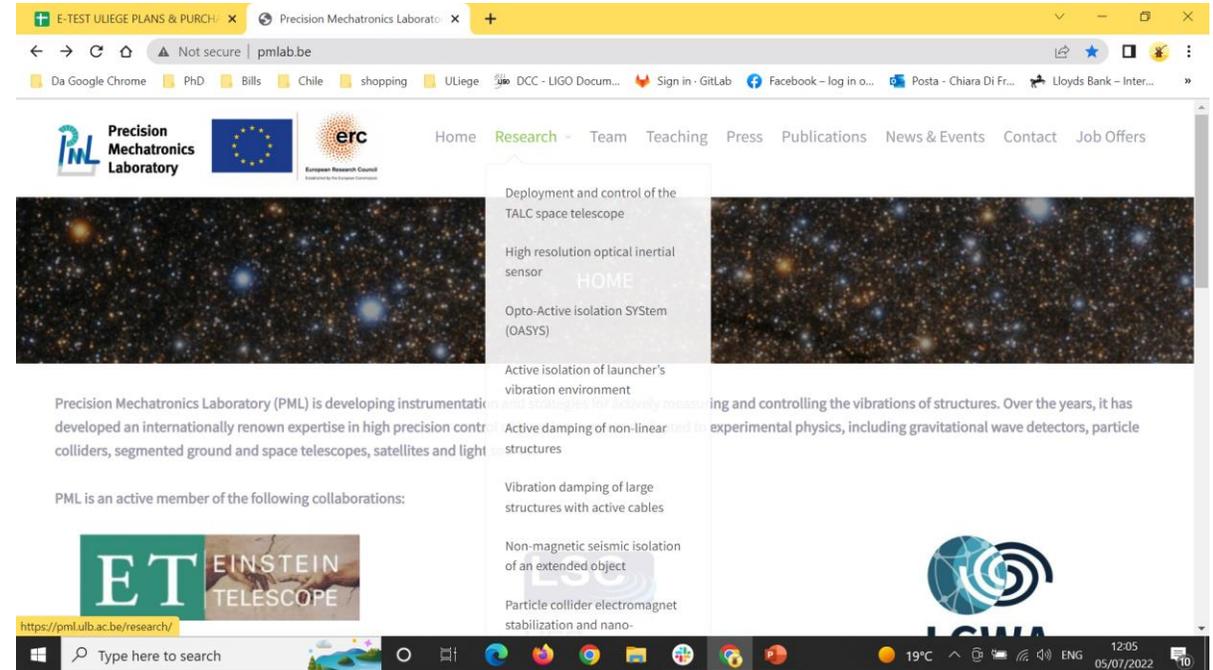
# A lab dedicated to seismic science

- There are several topics we cover, split between Université de Liège, Université Libre de Bruxelles and UCLouvain
- Main research is devoted to seismic isolation (for GW science but not only)
- We are a big group including 8 PhD students, 2 postdocs and several internships students from abroad, under the supervision of prof. Christophe Collette



# Vibration control for Geo- and Astro- observations

- Developing of new inertial sensors, sensitive at low frequencies
- Improving the sense and control of the gravity acceleration, in order to potentially reduce Newtonian noise
- Research on active control of vibration systems
- Testing the feasibility of the Einstein Telescope mirror isolation: E-TEST project at ULiege



# A bit about myself

- **PhD** at University of Birmingham (UK) on seismic isolation technologies (<https://etheses.bham.ac.uk//id/eprint/12495/>)
- **Postdoc** at Université de Liège (Belgium), where I'm responsible for the management of the E-TEST ERC project
- In collaboration with Nikhef and UoB, I'm publishing my work on laser stabilization for the 6D isolator ERC project
- Mentoring PhD students at ULiège
- Following-up my work on LIGO Hanford site on seismic isolating the active platforms via positions sensors connections
- I am an editor of the LIGO magazine, my last two articles are online
- CQG E-TEST recently published (*Ameer Sider et al 2023 Class. Quantum Grav.*)



**Fun stories from the LIGO sites**

### How do you travel the LIGO Arms?

How do you get from one side to the other of each arm. Traveling by car is the obvious solution, but there are some complications. I personally experienced the difficulties of driving a car to the LIGO sites. Definitely a funny experience, considering that usually such driving is a joy!

**Chiara Di Fronzo**

**How to survive a snake bite**

It happens that the best spots to go for from human living places are the ones that are the most dangerous. This is because a very specific environment is needed to be able to detect a gravitational wave or to test advanced technology.

**Chiara Di Fronzo**

**Essential skills for Researchers**

It happens that the best spots to go for from human living places are the ones that are the most dangerous. This is because a very specific environment is needed to be able to detect a gravitational wave or to test advanced technology.

**Chiara Di Fronzo**

**Snake Bite Kit**

**Information**

# The E-TEST project at ULiège

- Research at ULiège is ongoing to improve inertial sensors and to validate the advanced technology for new GW detectors, as Einstein Telescope (ET)
- At ULiège, the prototype E-TEST is under construction to test the hybrid technology mentioned before
- For more info about E-TEST in general, please visit:

<https://www.etest-emr.eu/>



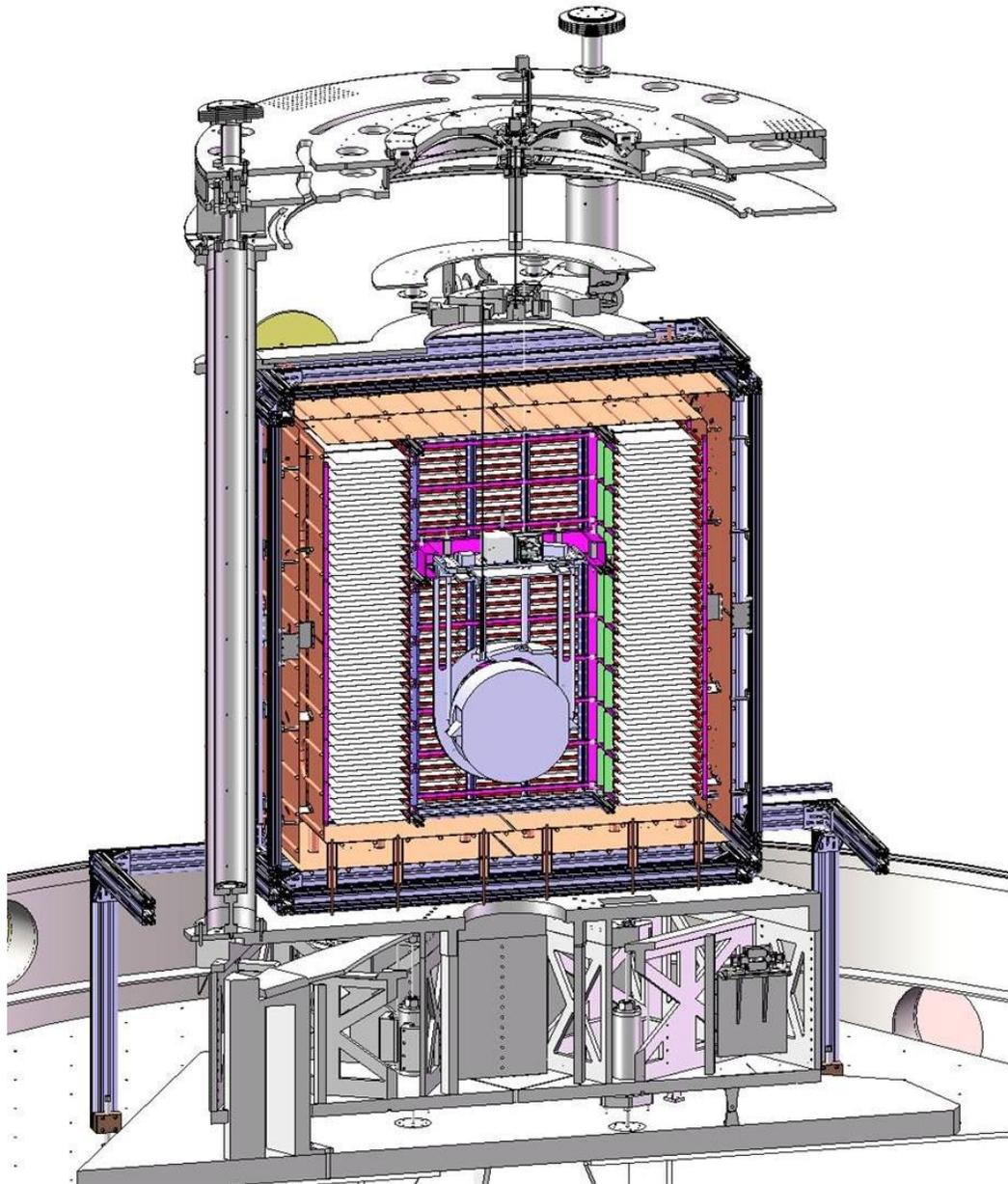
Work in progress.  
Credit: Haidar Lakkis

# E-TEST objectives

- Large mirror (100 Kg)
- Cryogenic temperature (10-20 K)
- Isolated at low frequency (0.1-10 Hz)
- Compact suspension (4.5 meters)

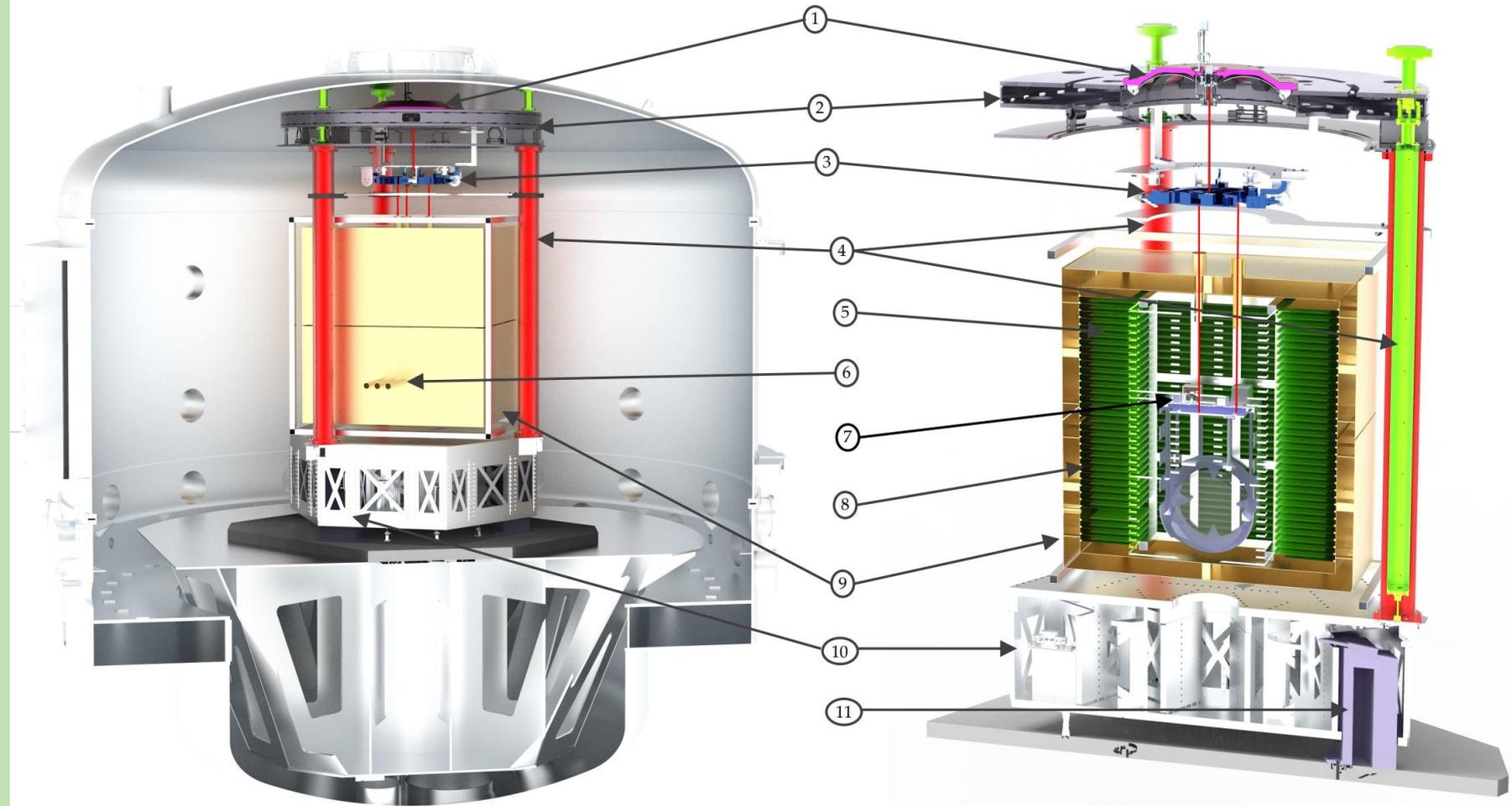
## E-TEST feasibility strategy

E-TEST is a project funded by the Interreg Euregio Meuse-Rhine and ET2SME consortium, which allow us to capitalize on existing infrastructure at Centre Spatial Liège (CSL) for the construction of the facility.



# E-TEST Isolation System - Conceptual Design

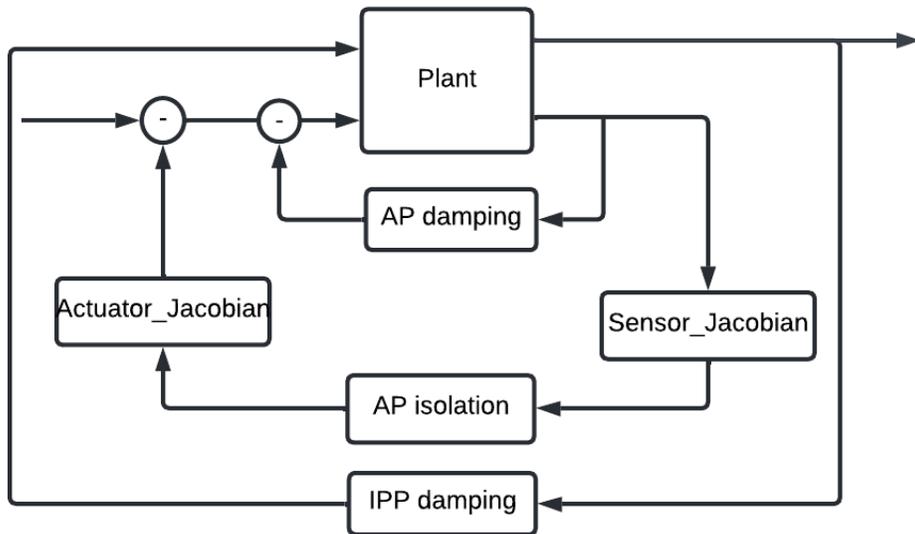
- 1) GAS filter.
- 2) Top stage.
- 3) Marionette.
- 4) Inverted pendulum legs within pipes that support a reference ring below the top stage.
- 5) Inner cryostat which has the interlacing fin type heat exchanger.
- 6) Three access points for outside experiments to interact with the cryogenic mirror.
- 7) The inner cryostat is attached to the cold platform.
- 8) Outer cryostat which provides a cold environment and houses the 100 kg silicon mirror.
- 9) Active platform.
- 10) Three large blades
- 11) Support pillar on the ground.



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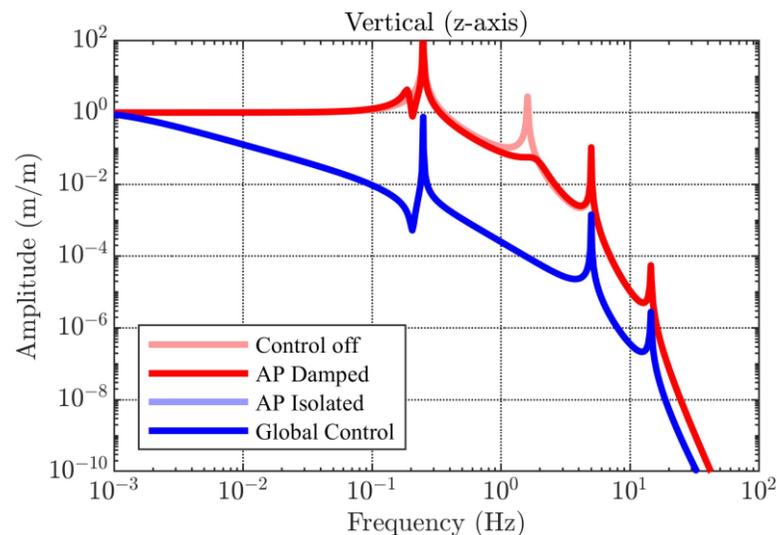
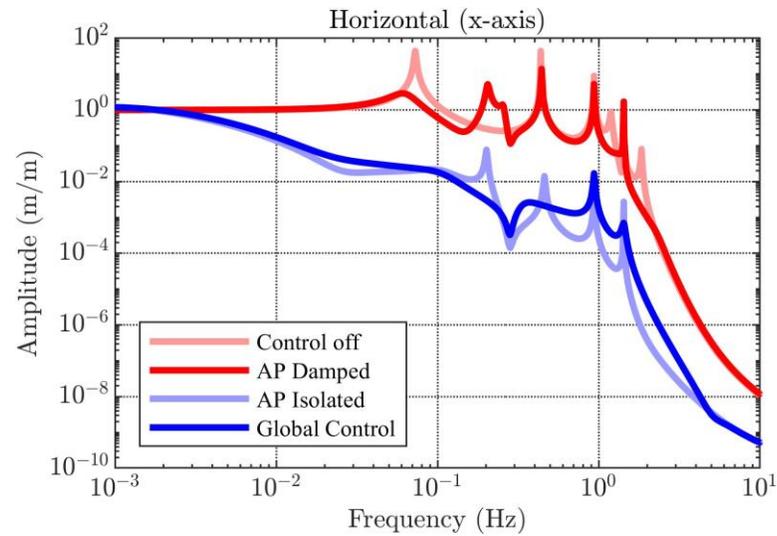
# Multi-body model and control strategy

## Schematic Control of E-TEST

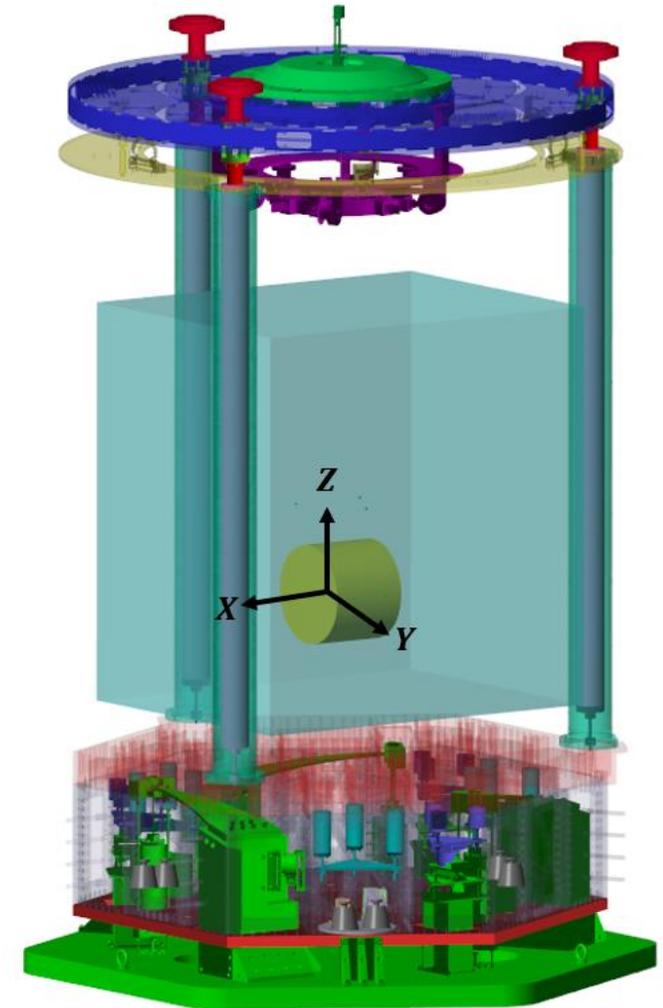


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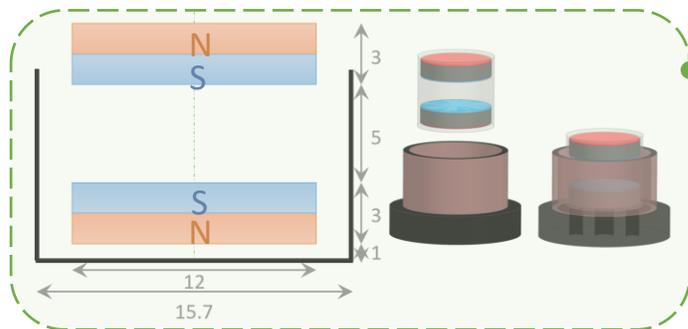
## Transmissibility (Mirror/ground)



## E-TEST Simscape model

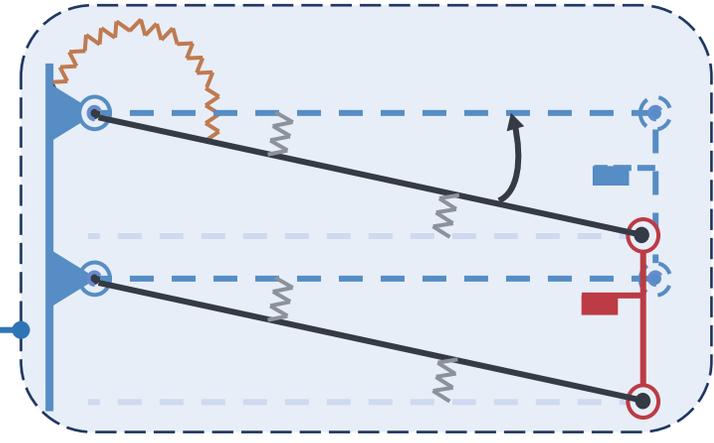
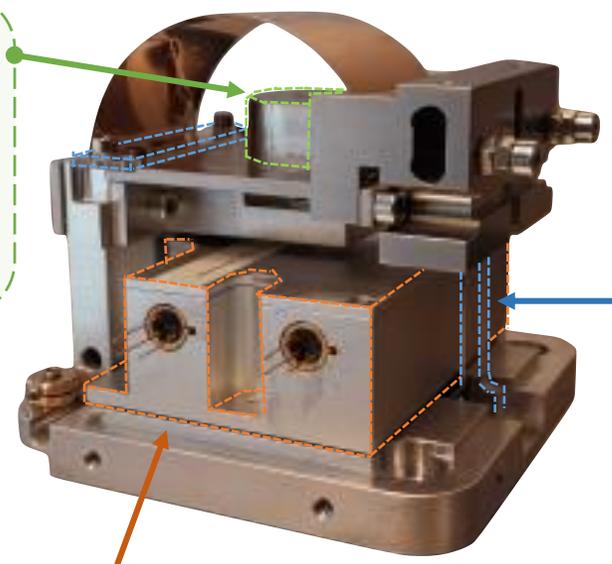


# Room T inertial sensor for the E-TEST project



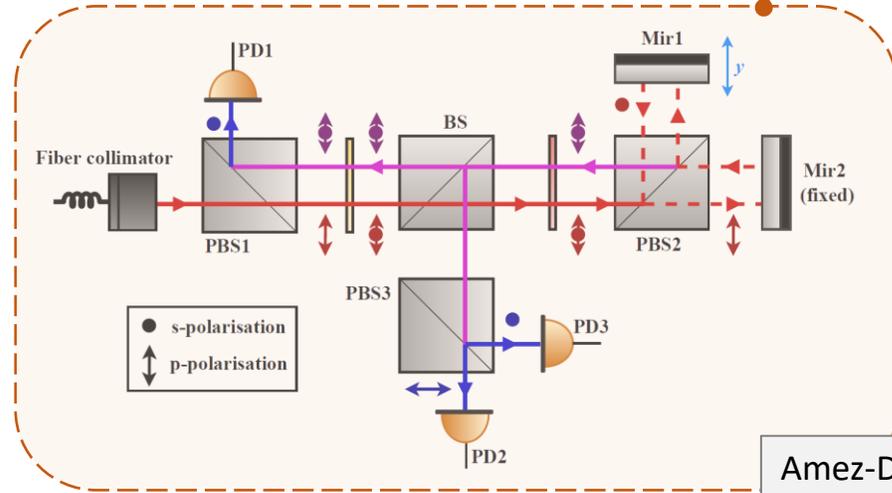
## Feedback actuator:

- Moving magnet VCA.
- Self-shielded quadrupole magnet.



## Mechanics

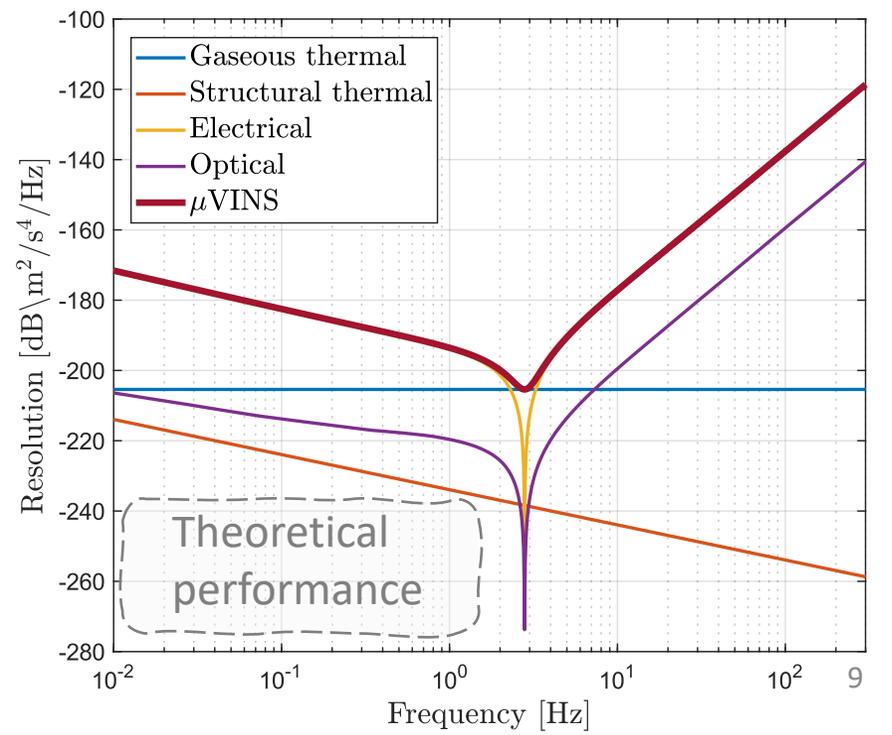
- Long period:  $f_0 = 2$  Hz
- Low loss fused-silica flexures.
- Linear mechanical guide for the optical readout.

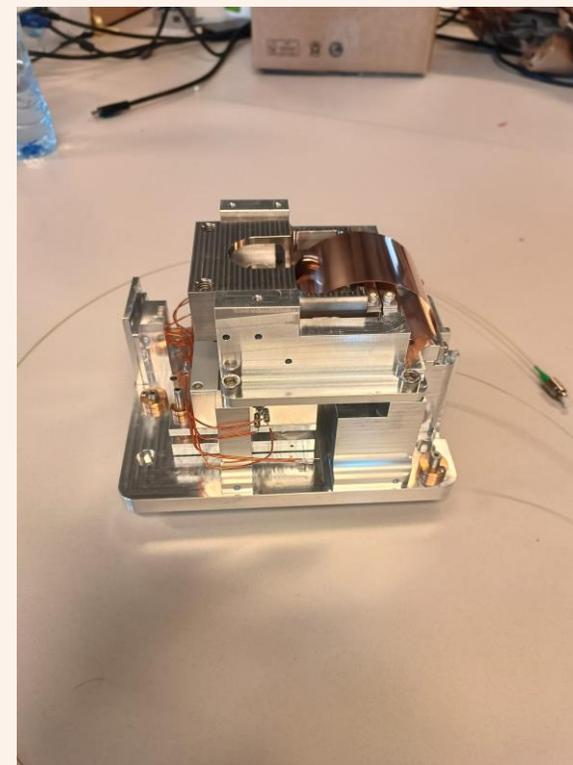
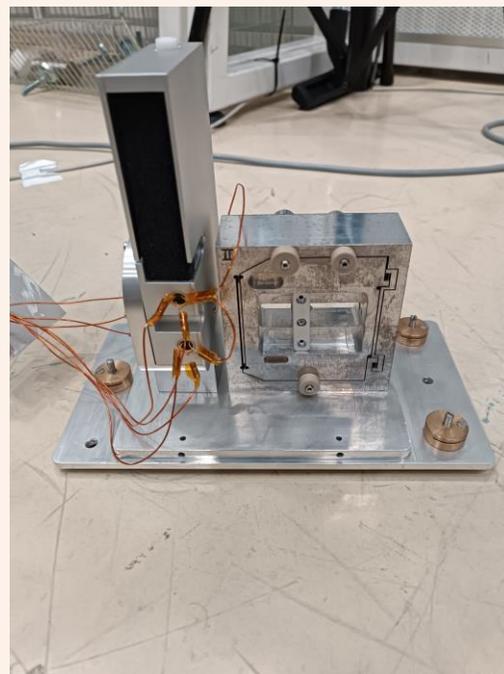
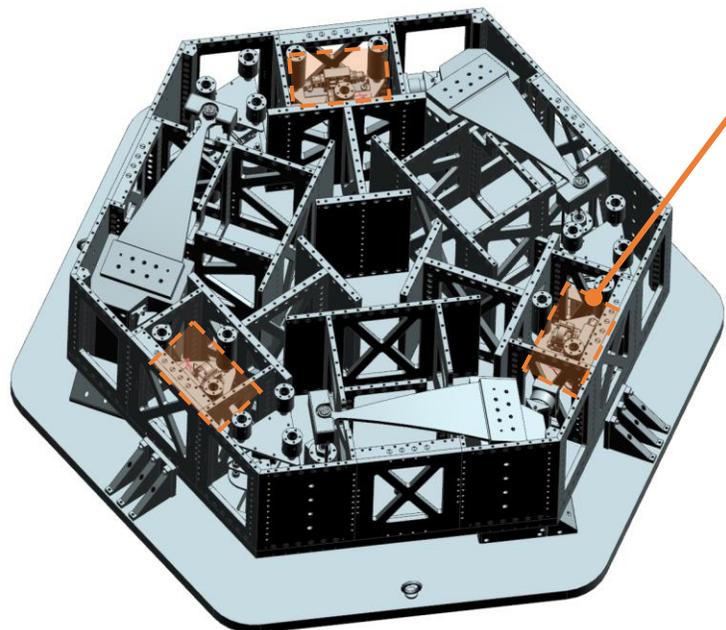
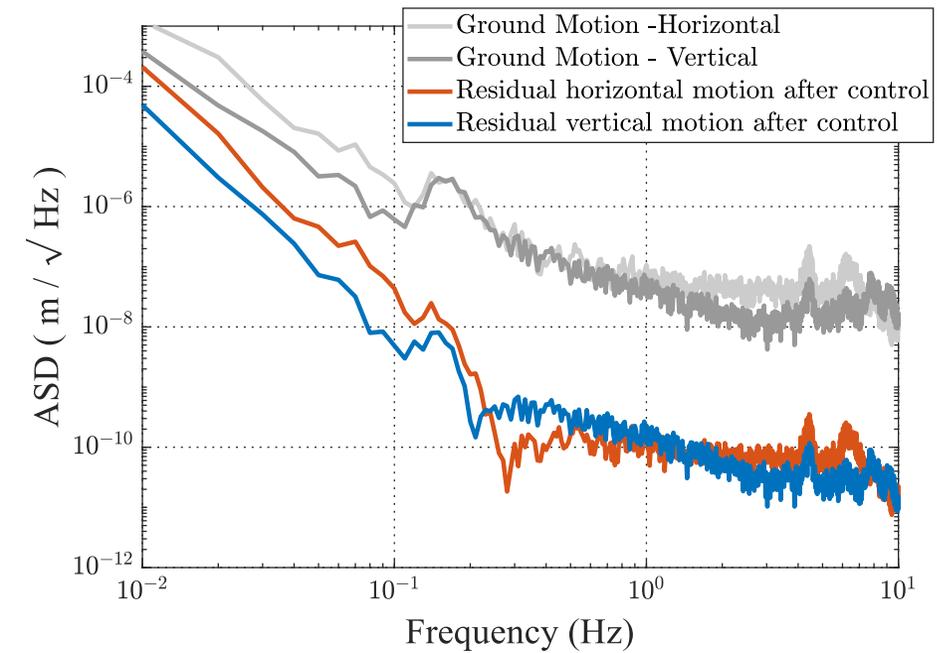


## Quadrature Michelson interferometer

- High dynamic range.
- High resolution:  $2 \times 10^{-13}$  m/VHz.

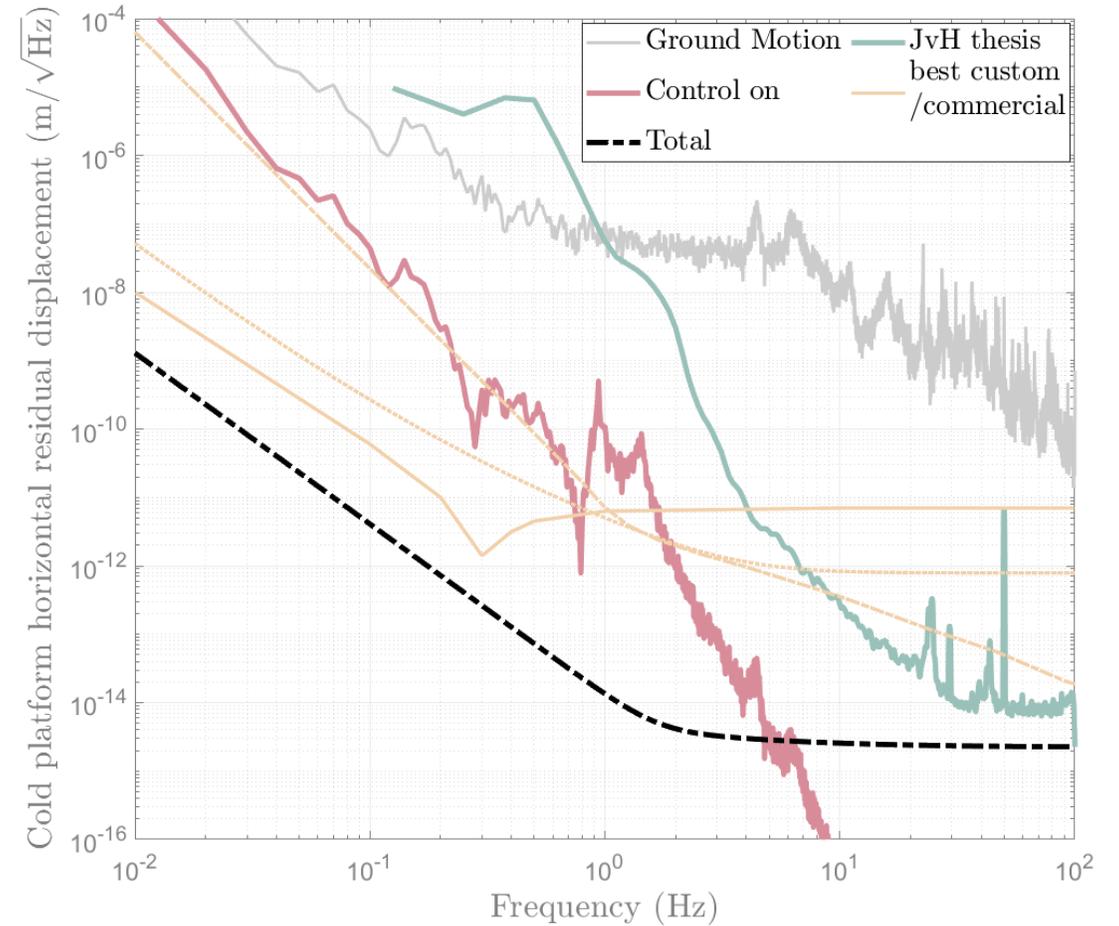
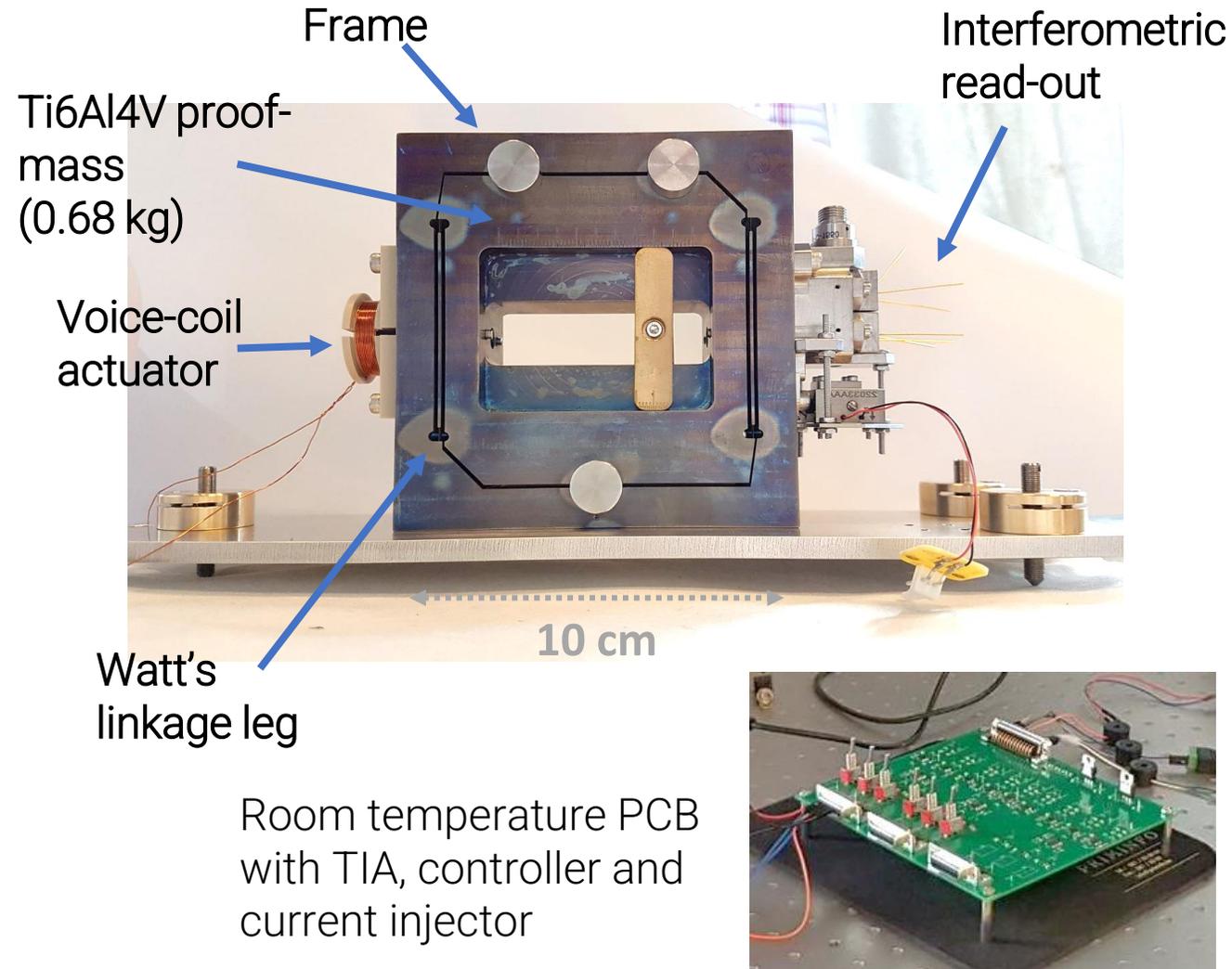
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# Horizontal cryogenic inertial sensor for E-TEST

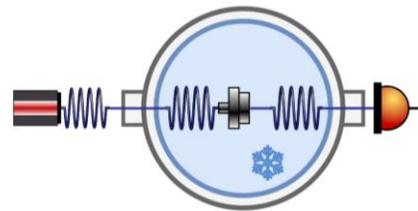
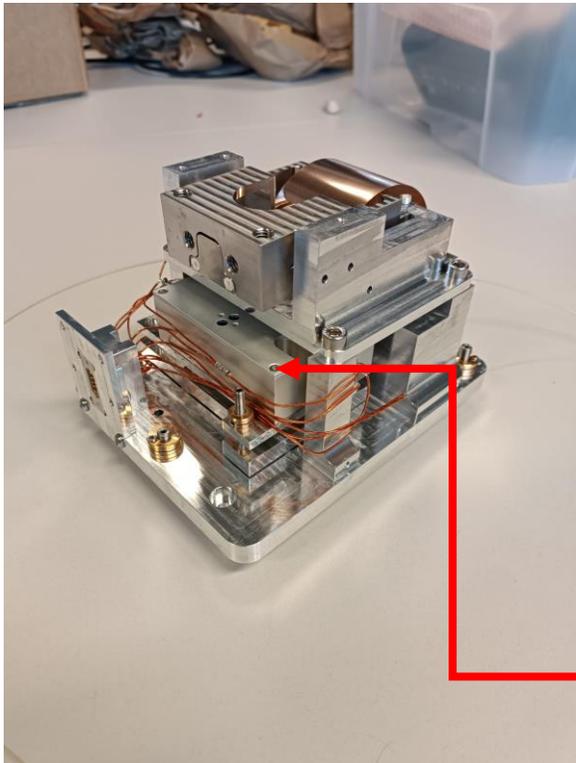


Credit: Morgane Zeoli

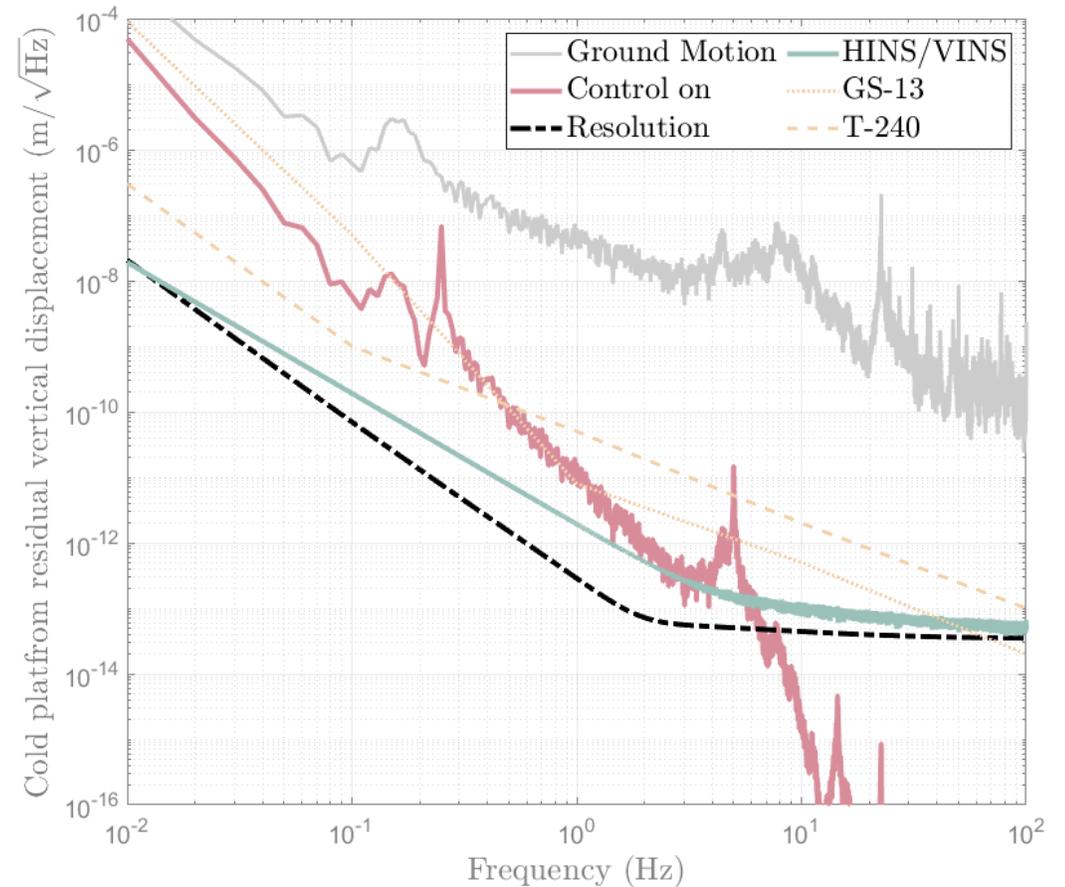
# Vertical Cryogenic inertial sensor for E-TEST

E-VINS design adapted for cryogenic working conditions

A test campaign was taken in collaboration with RWTH Aachen to select the optical elements that works the best in cryogenic conditions (collimators, photodiodes, polarization, alignment, etc). The results are used for both CSIS-V and H.



Interferometric read-out

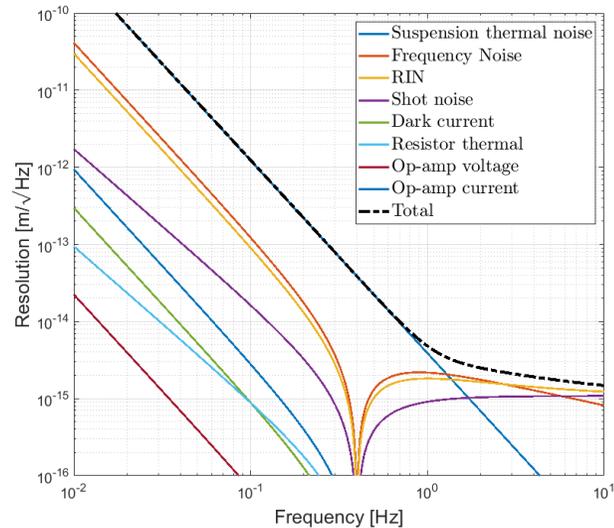


Credit: Morgane Zeoli

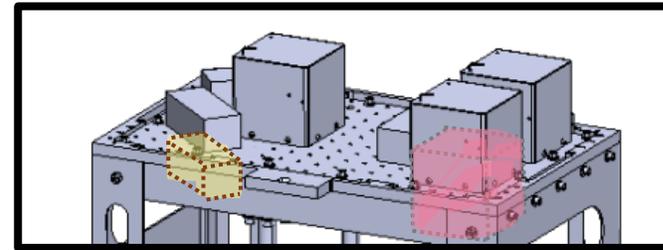
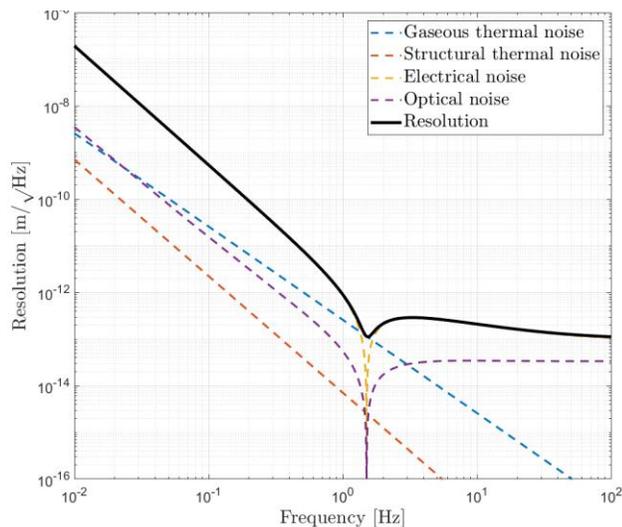
# Final use and test bench

- Monitor residual motion with fm/ $\sqrt{\text{Hz}}$  sensitivity down to 1 Hz
- Exploit E-TEST cold platform low-vibration cryogenic environment to perform a self-noise measurement

Horizontal



Vertical

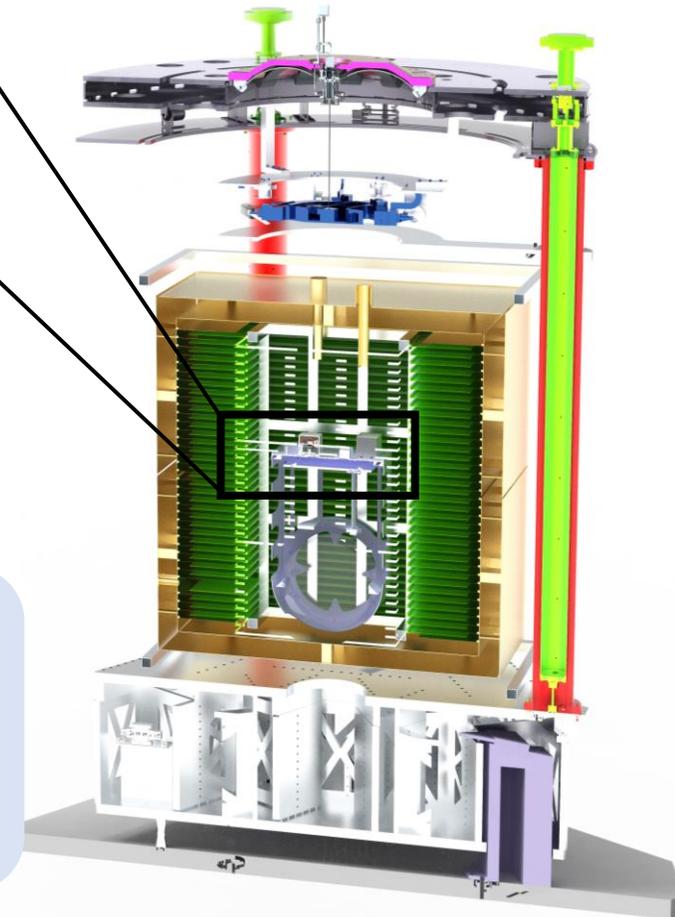


After the CSIS validation with E-TEST, an adaptation of this sensor will be embedded in the **LGWA (Lunar Gravitational Wave Antenna)**

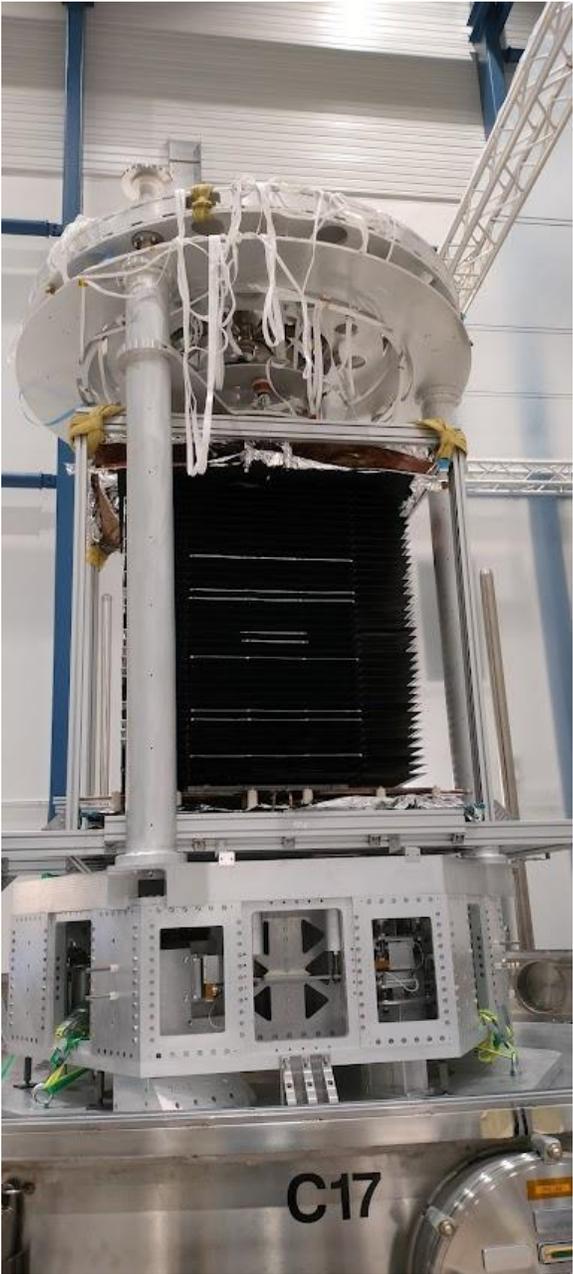


## GOAL:

- Measure the GW passing by the Moon
- New GW spectrum
  - Selene physics



# Current status of E-TEST





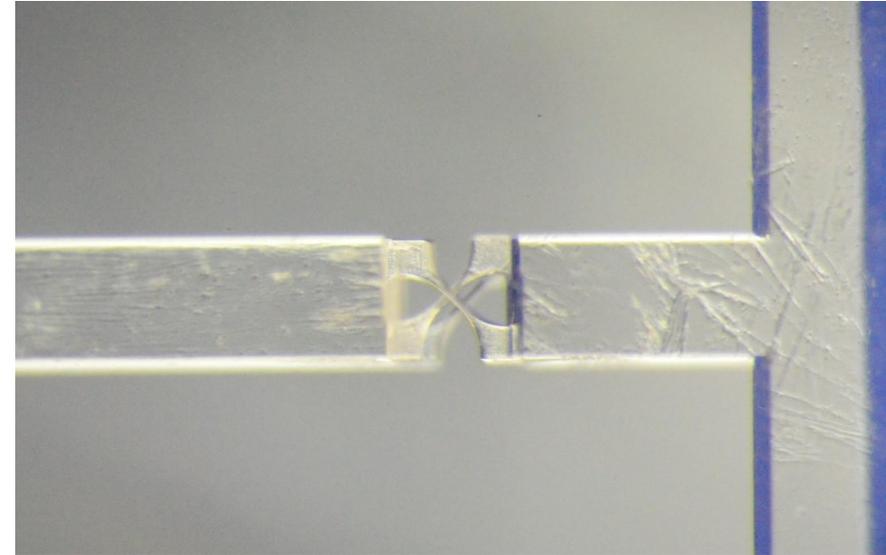
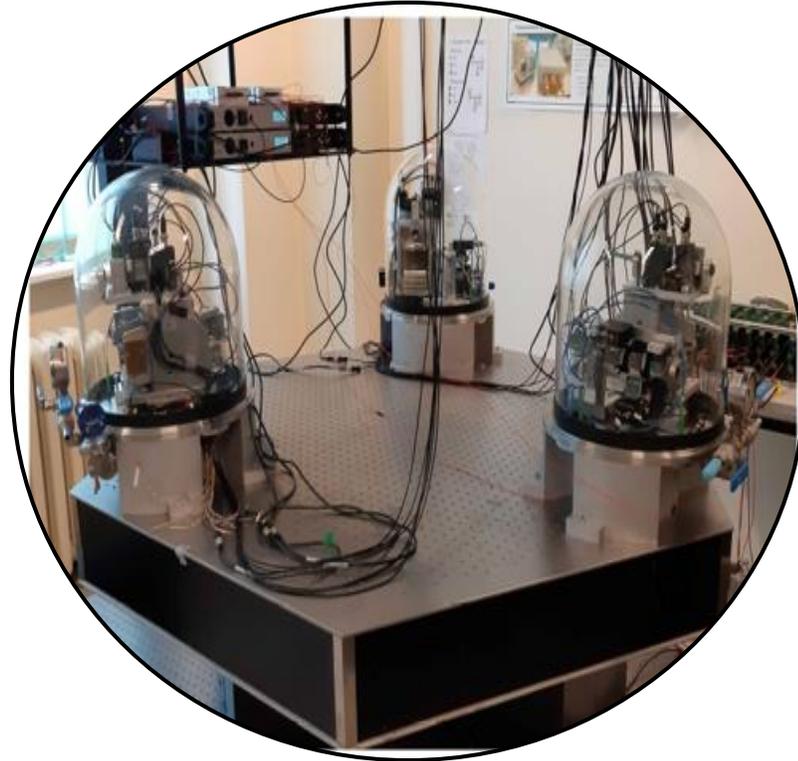
Vacuating and cooling started and it's going to take 3 weeks (N done, He started.  $T = 90\text{ K}$ ).

During that time, we monitor the sensors, apply control and test the responses.

After that, we will dismount the cryostat and move E-TEST to another lab for suspension and seismic tests.

End 2024: installation of the silicon mirror and rods.

# Other than E-TEST



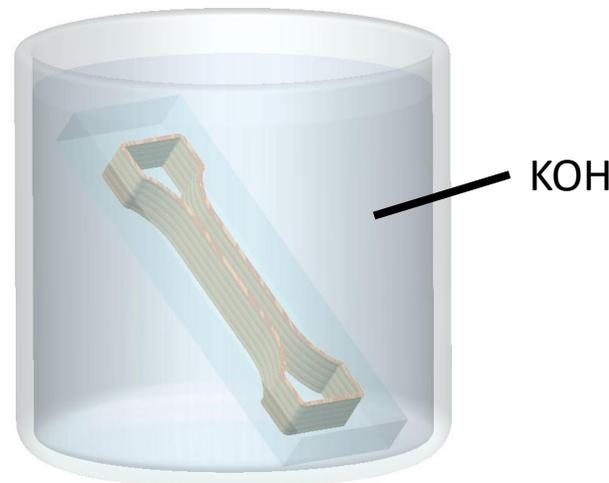
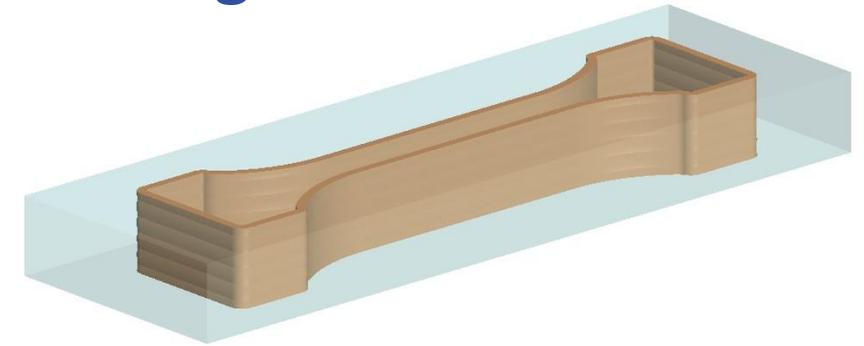
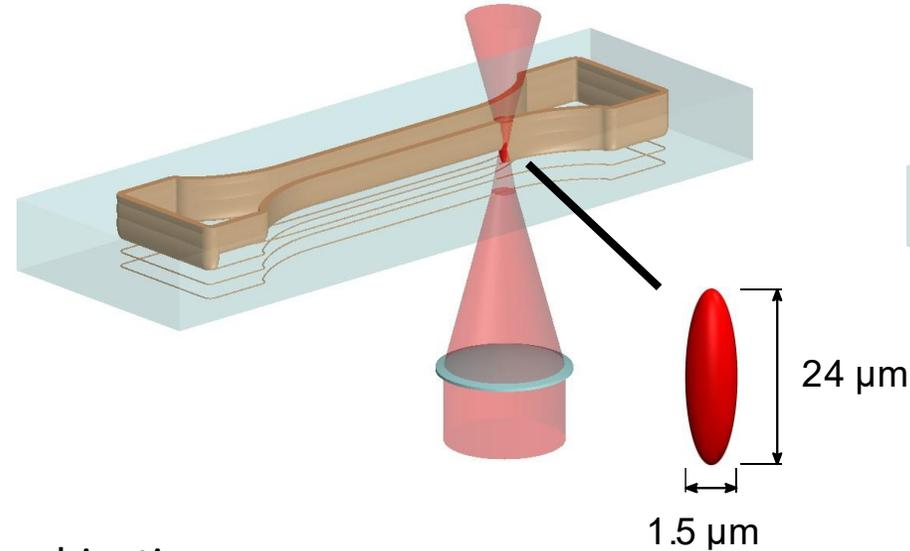
# Glass structure manufacturing: laser assisted etching



Laser objective



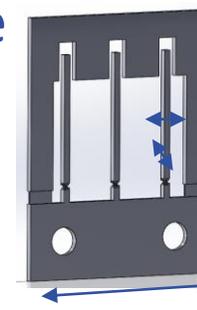
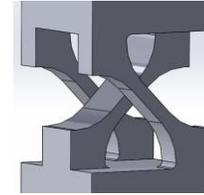
Substrate



- The laser path is defined to expose the shape of the desired part in a glass substrate
- Then, the substrate is placed in a KOH bath. Laser-exposed glass is etched 100x faster than non-exposed glass

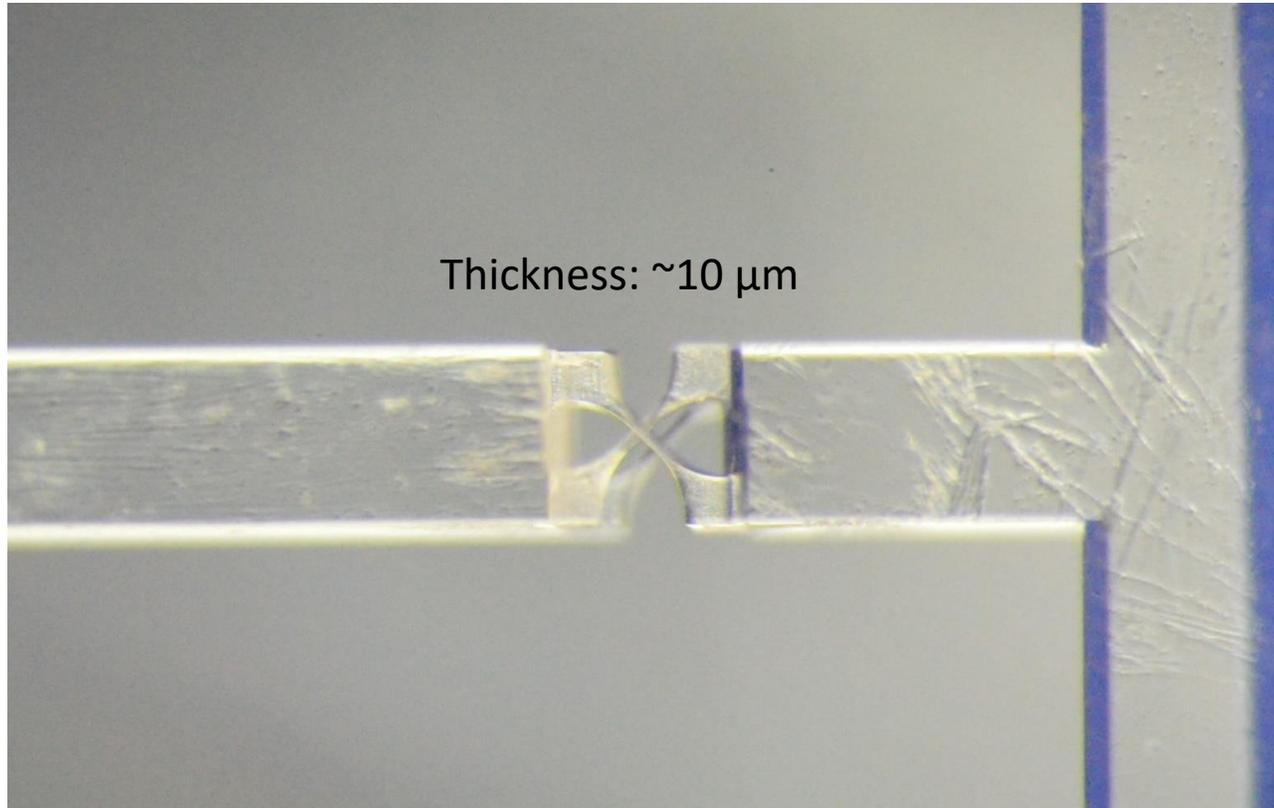


## Cross-spring pivot hinge



0.5 mm

10 mm

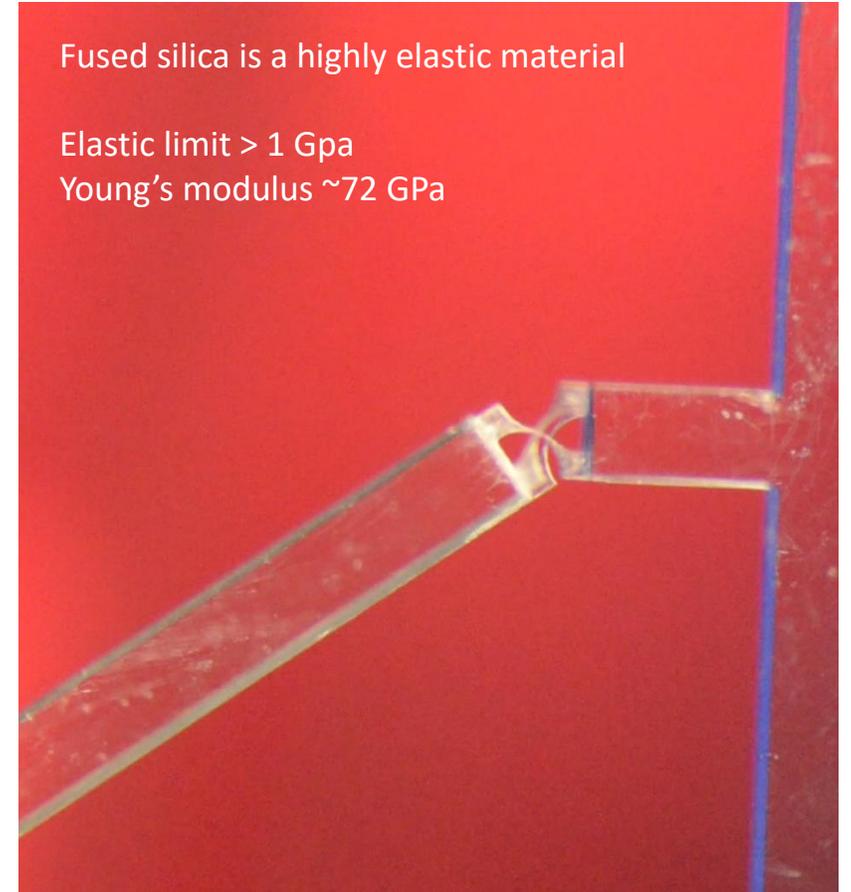


Thickness: ~10  $\mu\text{m}$

Fused silica is a highly elastic material

Elastic limit > 1 Gpa

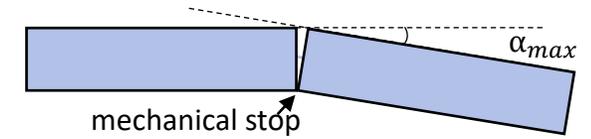
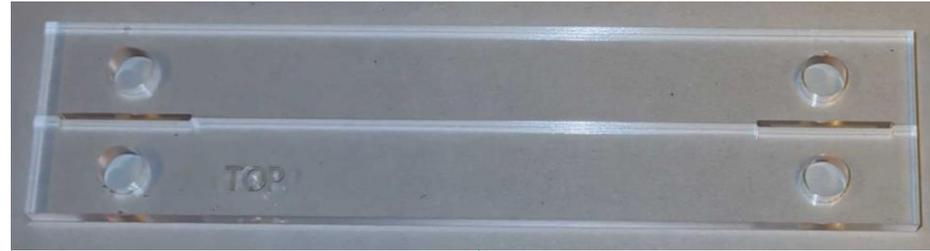
Young's modulus ~72 GPa



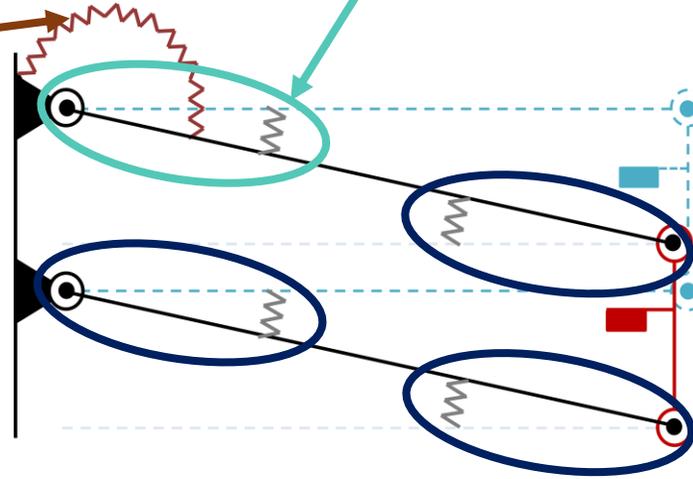
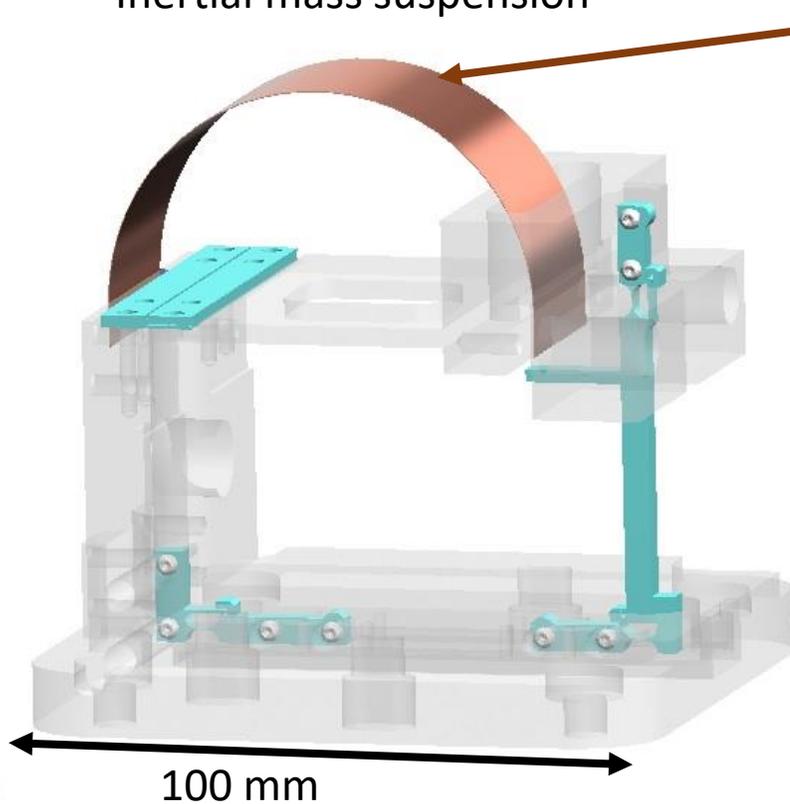
Thanks to the high aspect ratio of the manufacturing process, we are able to produce monolithic compliant mechanisms

# Inertial sensor $\mu$ VINS

Inertial mass leaf-spring hinge

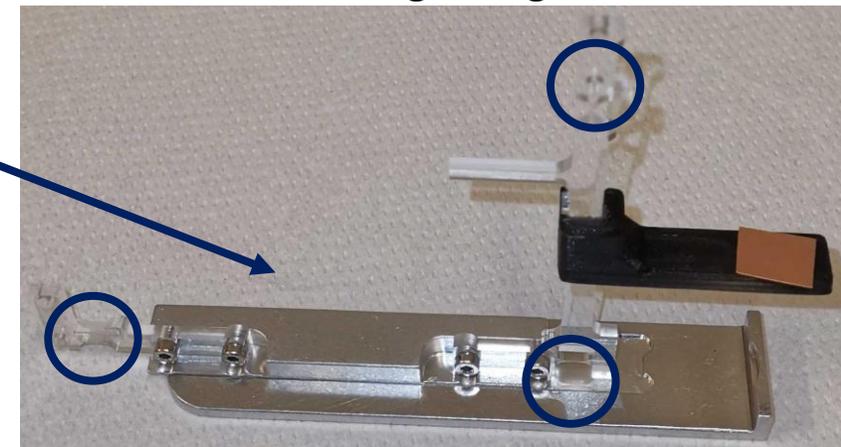


Inertial mass suspension

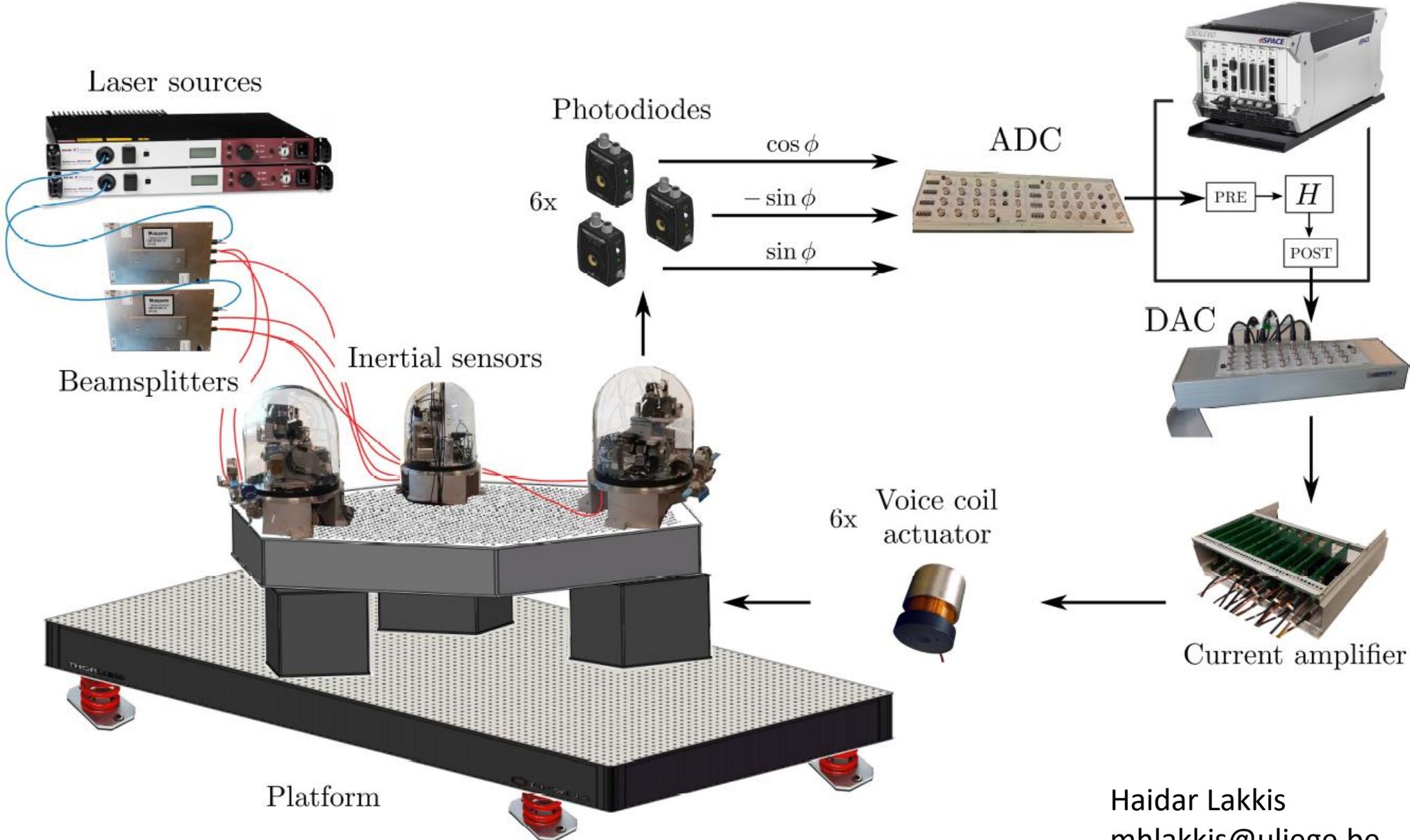


- Glass can be used for compliant joints to reduce drift and thermal noise (depending on the stress ratio between the joint and the suspension)

Mirror guiding

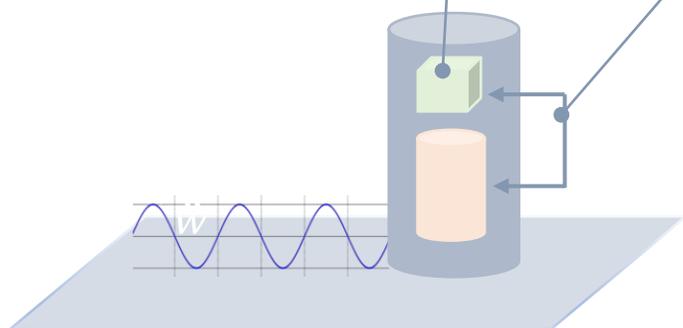
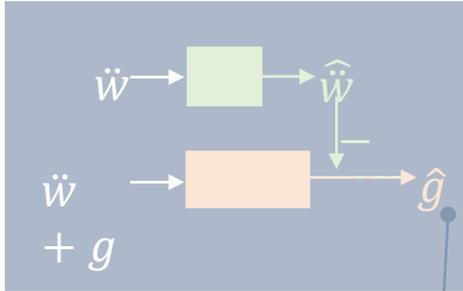


# Low-frequency isolation of 6DOF systems using inertial sensors

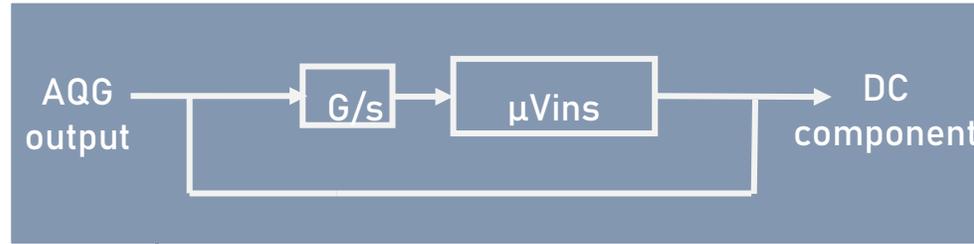


# Atomic Quantum Gravimeter ibridization

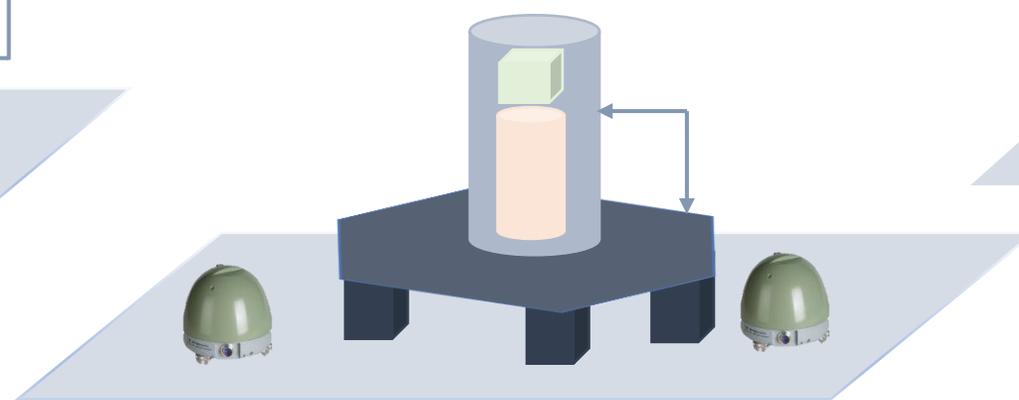
## 1. Improve the gravimeter



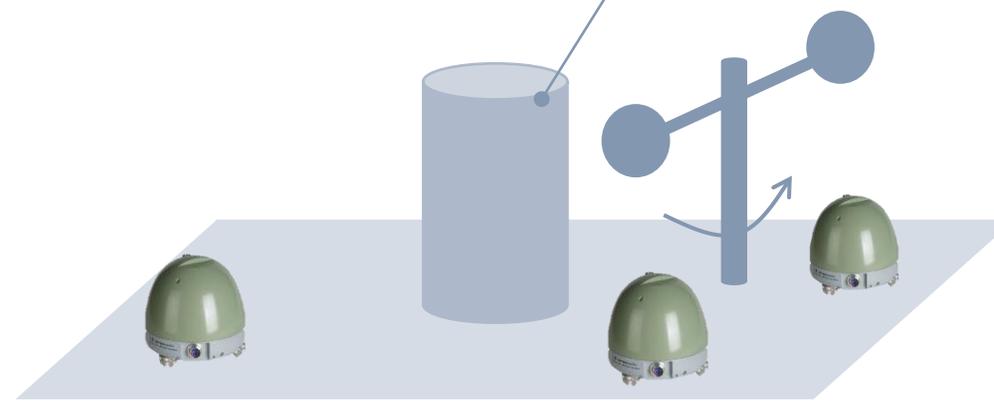
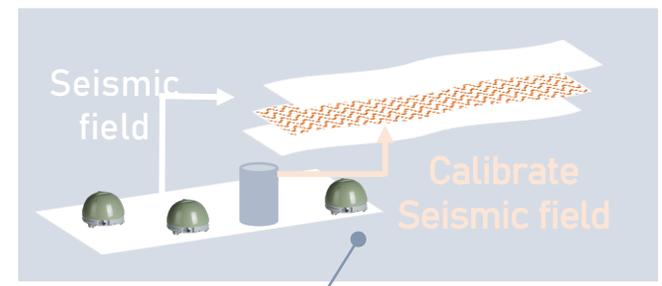
## 2. Double hybridization



## 3. Active Isolation



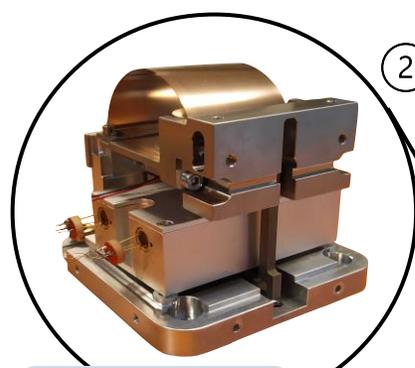
## 4. Seismic Newtonian Noise Calibration



# Overview

② Hybrid sensor: AQG + classical accelerometer

Broad bandwidth  
Continuous  
Long-term stability



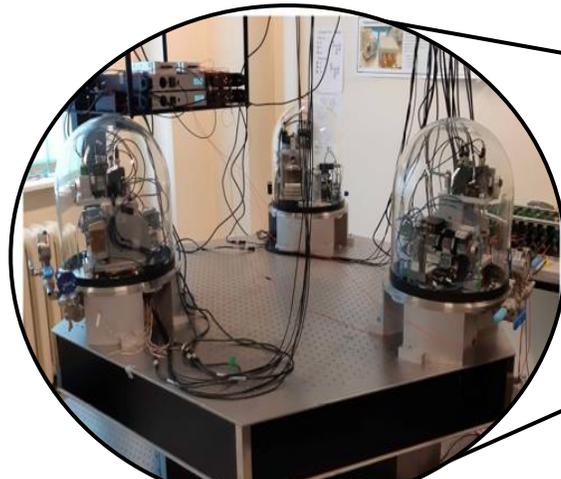
①  $\mu$ Vins



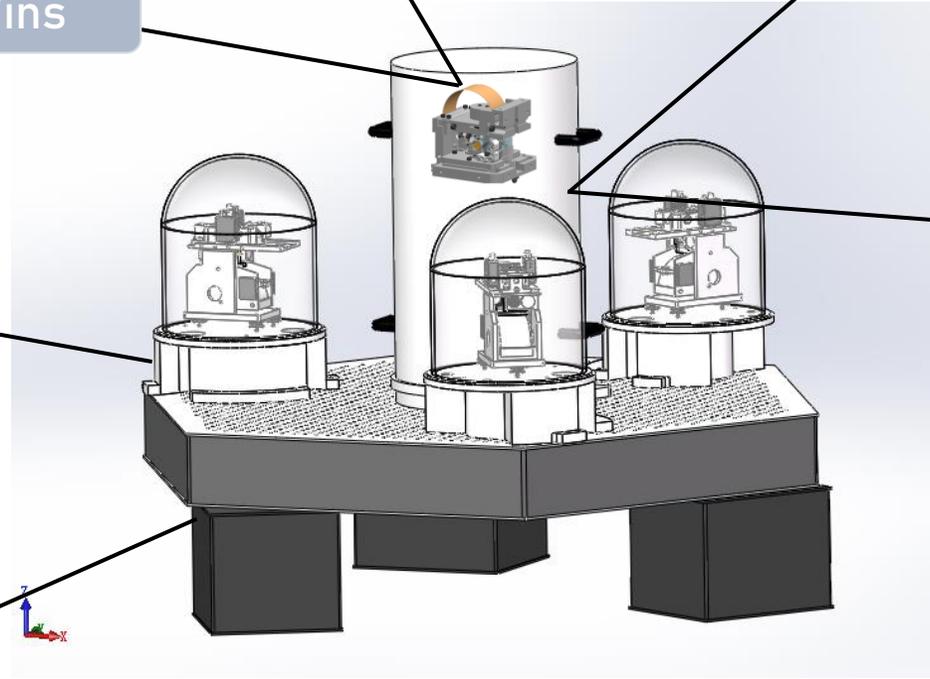
exail

Atomic  
Quantum  
Gravimeter

④ Newtonian  
Noise



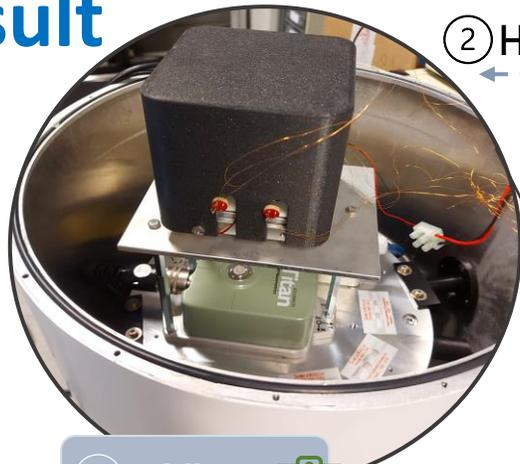
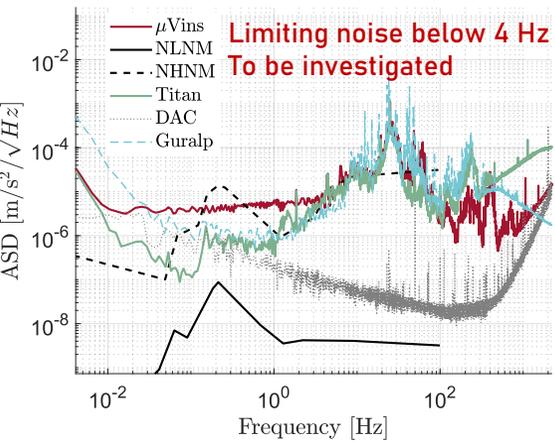
③ Active  
platform



Active isolation

Mayana Teloi  
mayana.teloi@uliege.be

# Preliminary result



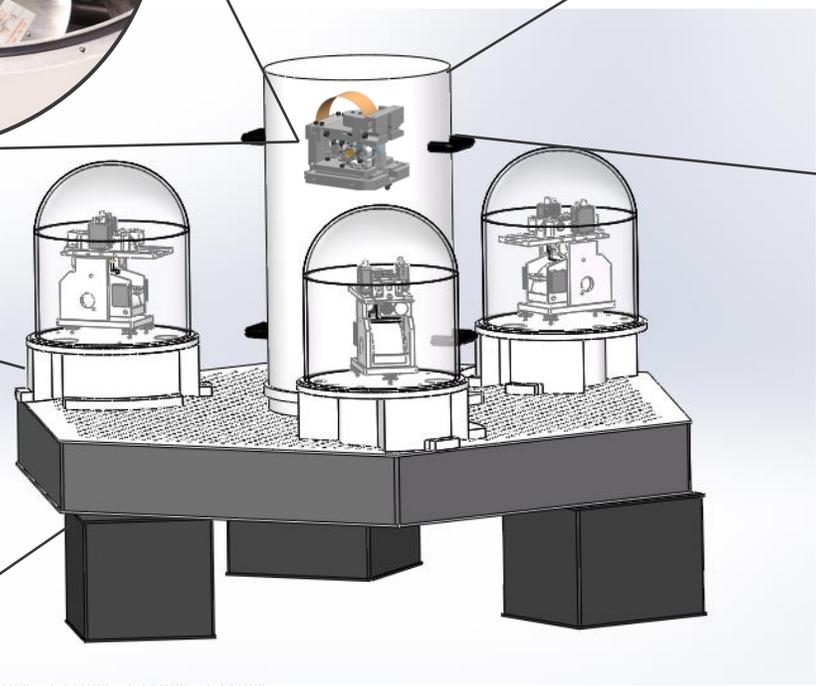
② Hybrid sensor: AQG + classical accelerometer



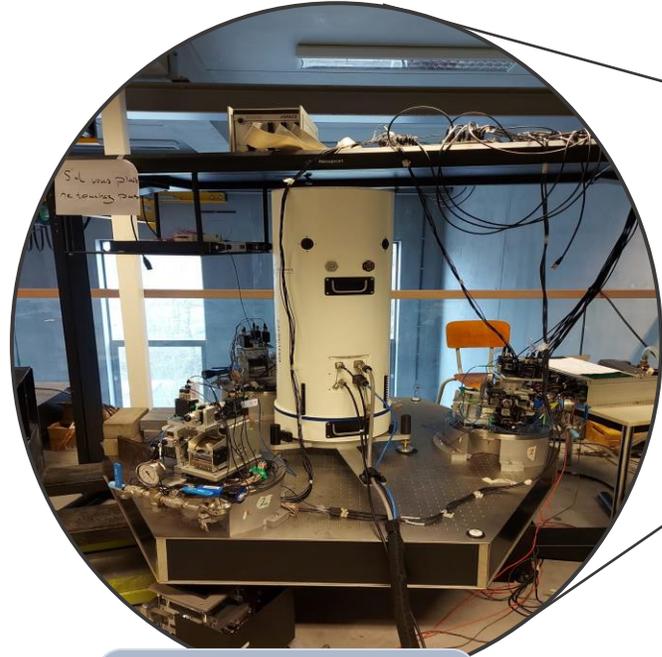
Atomic Quantum Gravimeter

Performance limited by ground vibrations in noisy environment

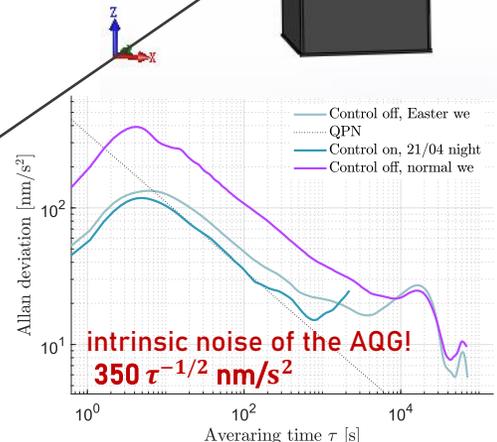
① μVins



④ Newtonian Noise



③ Active platform



With the active isolation, we are reaching the intrinsic noise of the gravimeter which is set by the laser. So ground noise is now completely removed from gravity measurement.

The gravimeter performance are limited by the ground cancellation strategy.

# General contacts and uni network

- Head of the group: Prof. Christophe Collette [christophe.collette@uliege.be](mailto:christophe.collette@uliege.be)
- Postdocs:
  - Dr. Chiara Di Fronzo [cdifronzo@uliege.be](mailto:cdifronzo@uliege.be)
  - Dr. Rasa Jamshidi [R.Jamshidi@uliege.be](mailto:R.Jamshidi@uliege.be)
- PhD students ULiege and ULB
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## Useful links:

### TDR

<https://arxiv.org/abs/2212.10083>

### E-TEST Project website

<https://www.etest-emr.eu/>

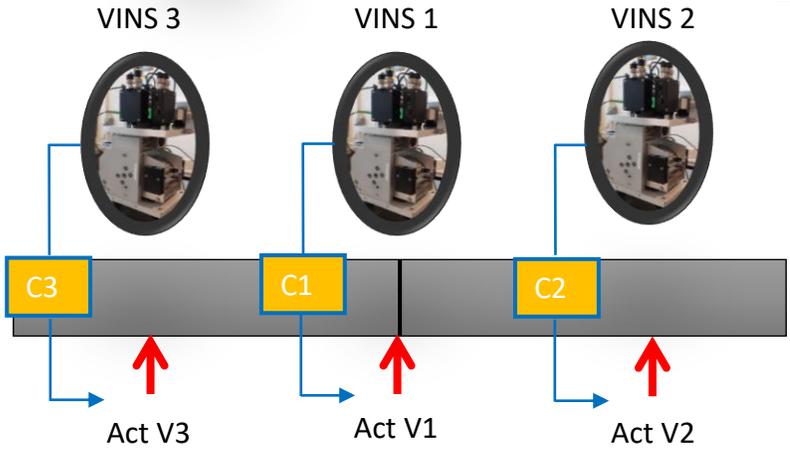
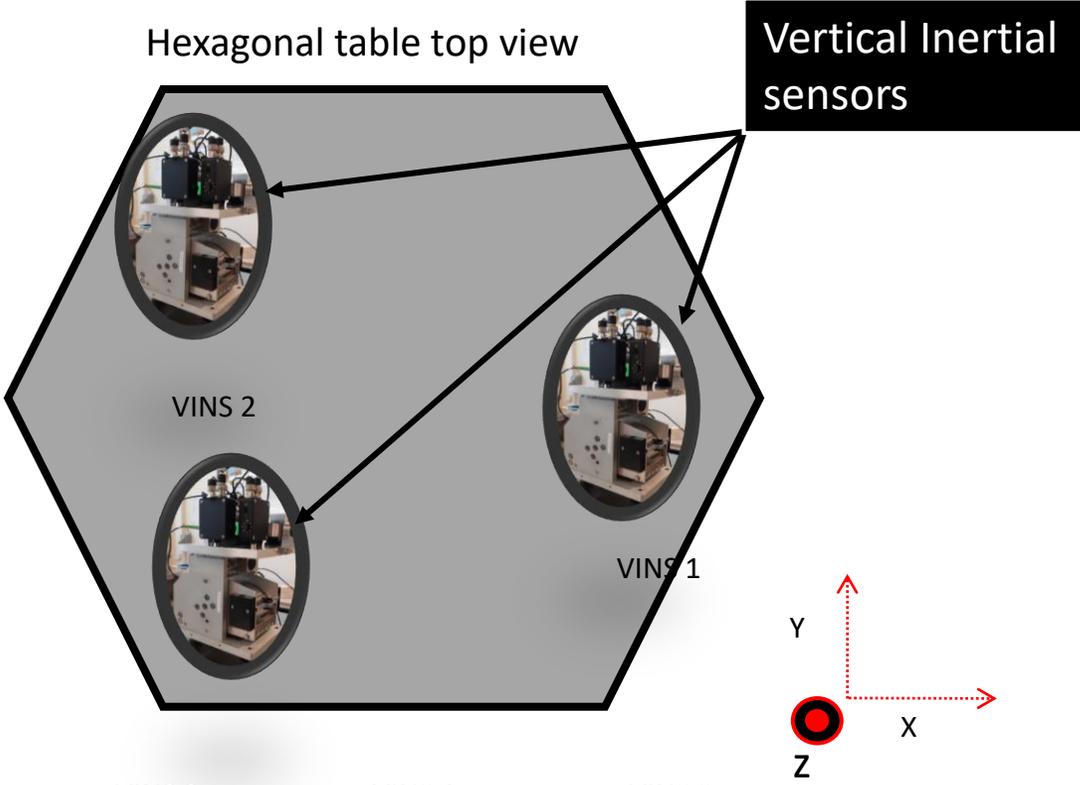
### PML website

<http://www.pmlab.be/>

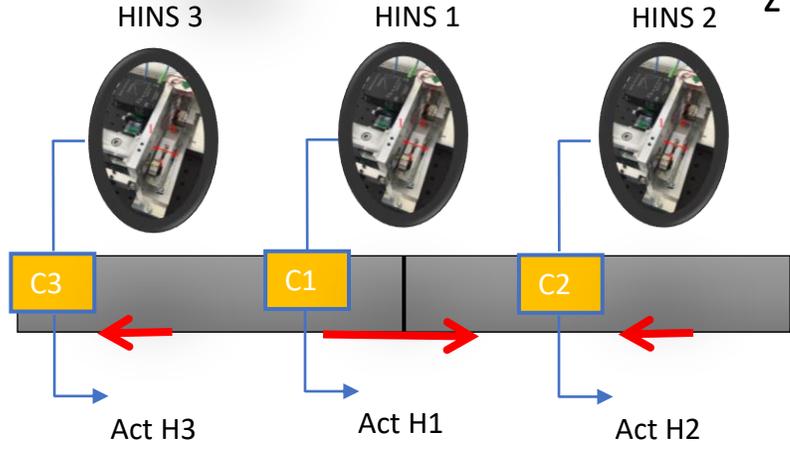
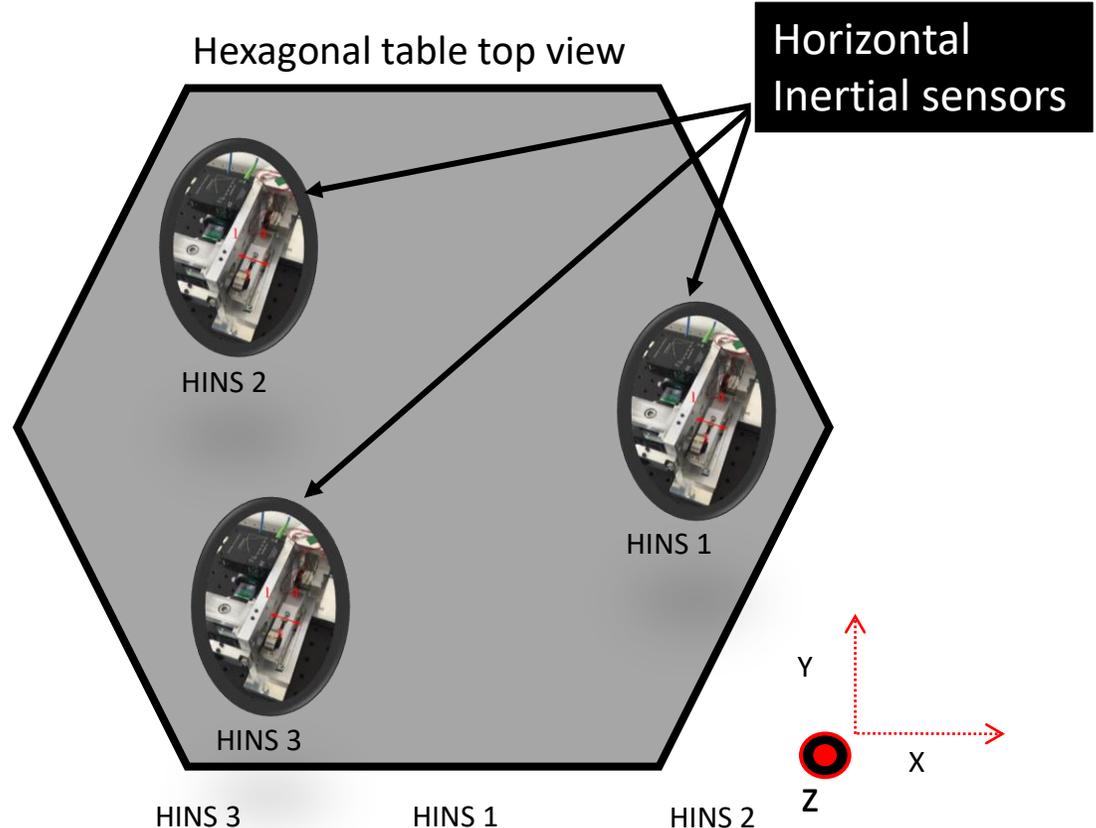


Back up

Decentralized Control :



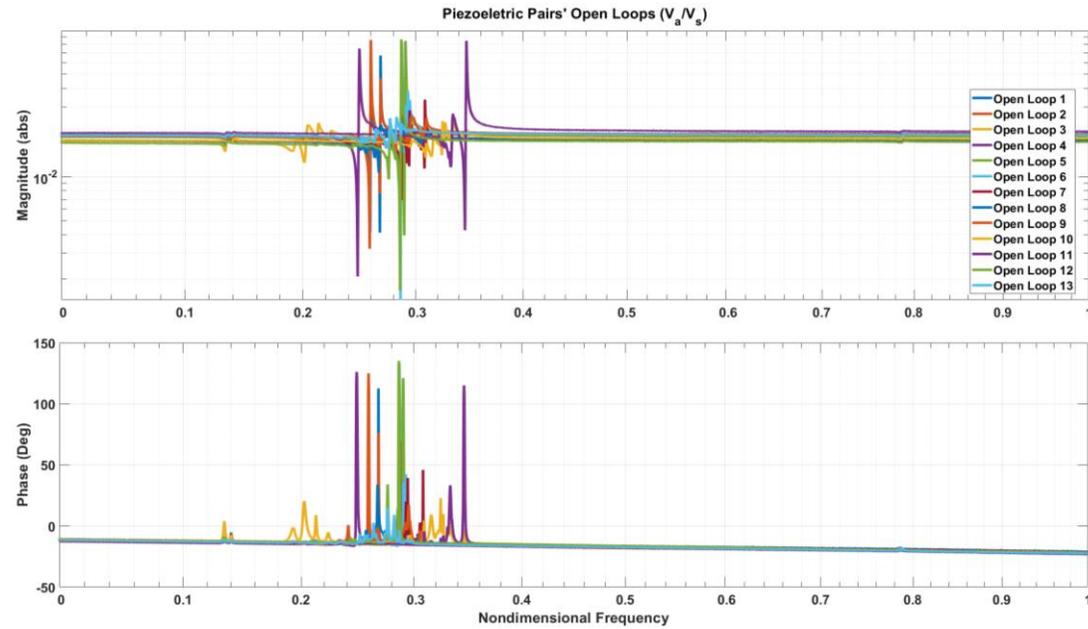
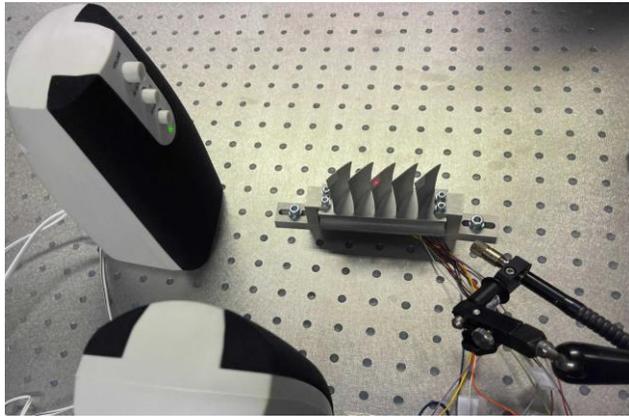
Vertical loops closed in a decentralized way



Horizontal loops closed in a decentralized way

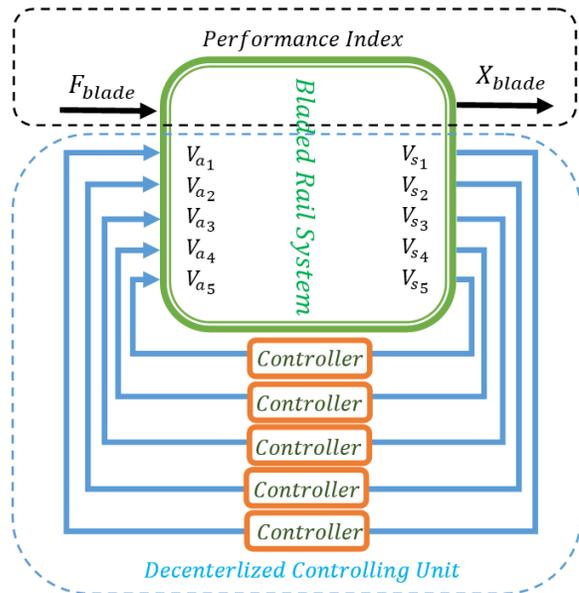
# Maverick experiment

## Bladed Rail

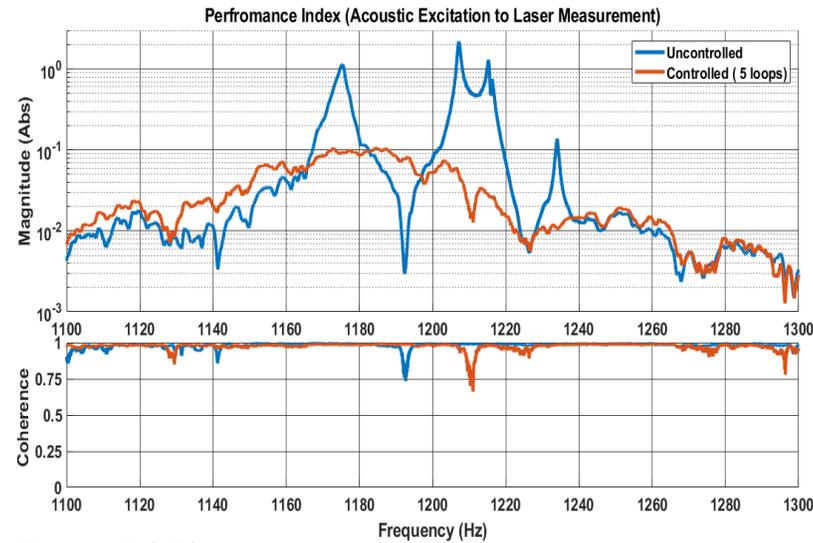


Open Loops of BLUM

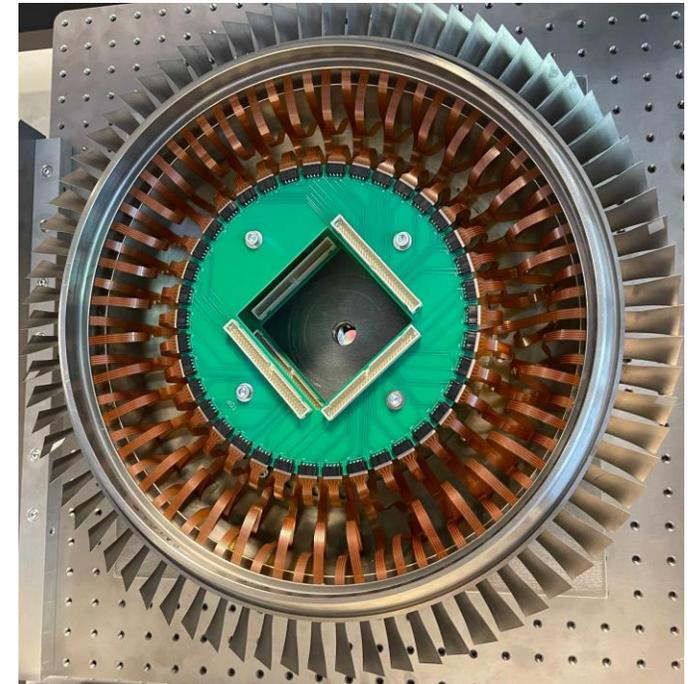
## Open Loops of Bladed Rail



## Performance Index of Bladed Rail

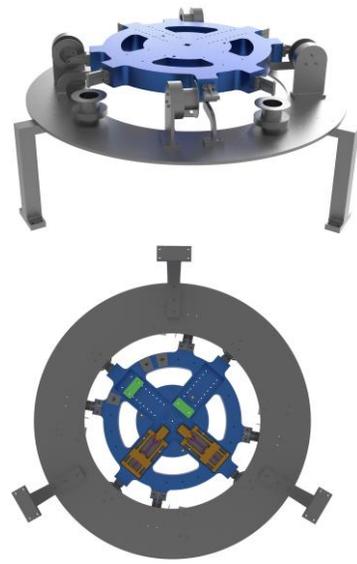


## BLUM



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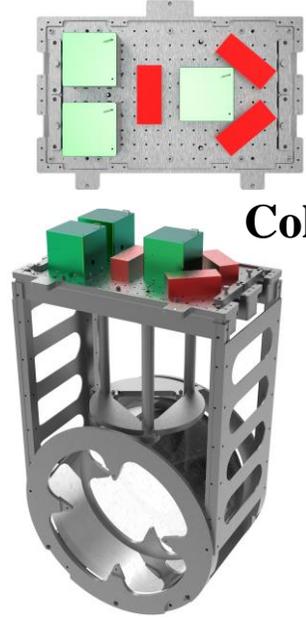
**Marionette**



**Cryostat**



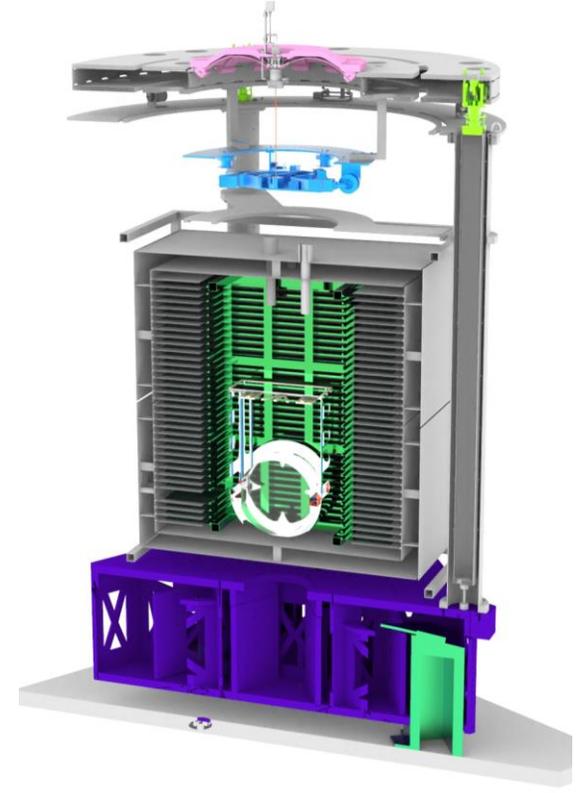
**Cold platform  
&  
Mirror**



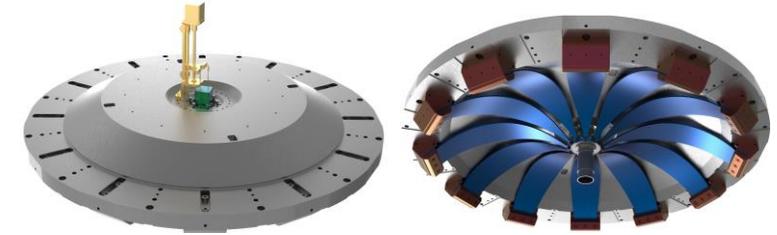
**Inverted Pendulum**



**Active platform**



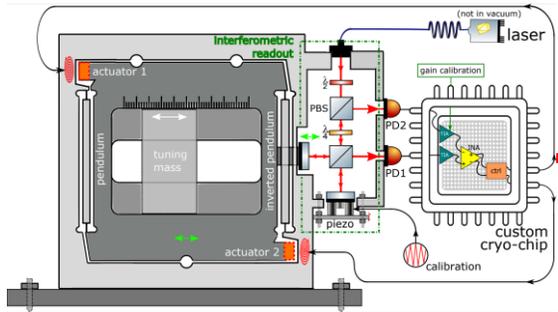
**GAS filter**



**Ameer Sider (PML)**  
**asider@uliege.be**

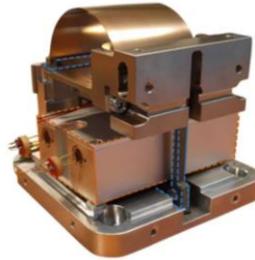
# Sensors & Actuators (vacuum compatible)

6 Cryogenic inertial sensors



Temperature sensors

6 Inertial sensors

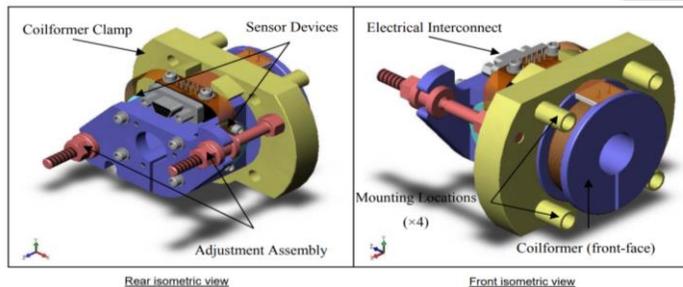


6 voice coil Actuators

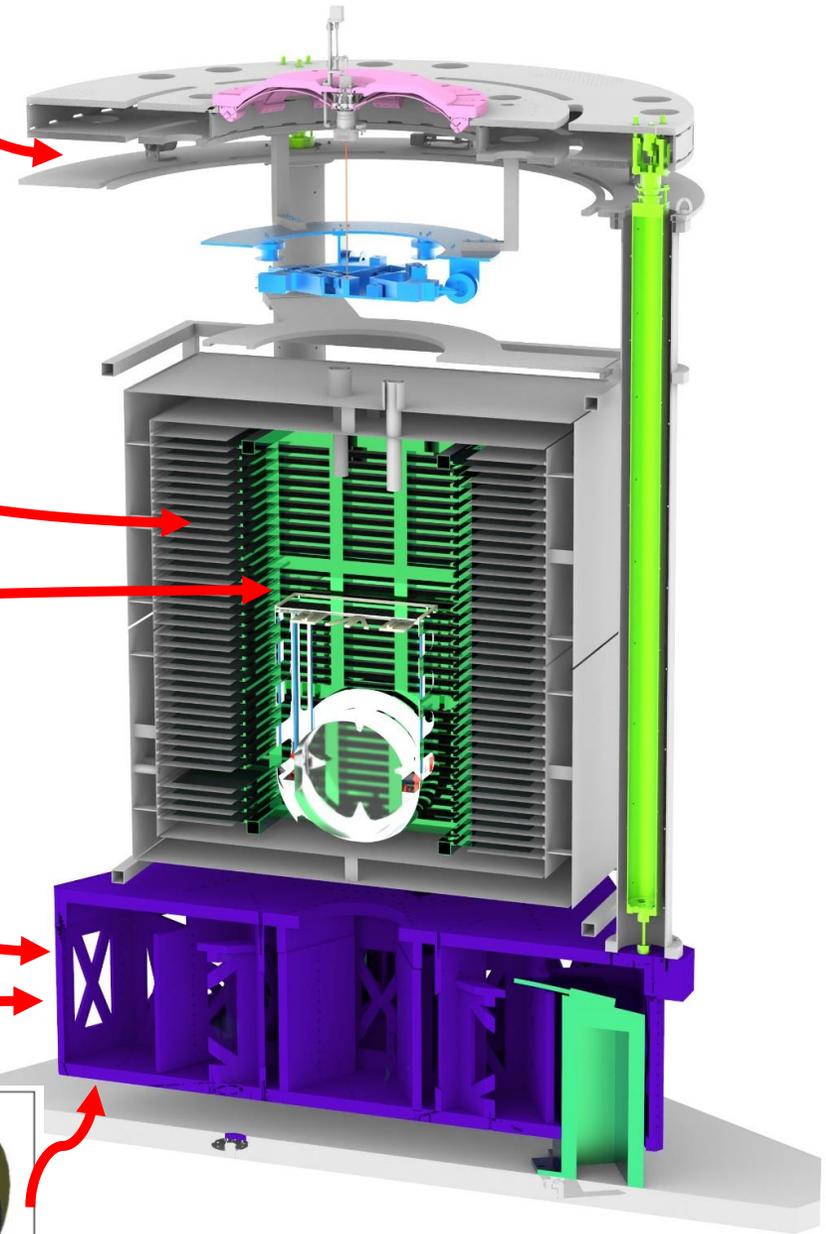


4 Optical Sensors

4 Voice coil actuators



6 BOSEMs



# Cross-spring pivot hinge

